

DESY Theory Workshop

Sept, 28th – Oct, 1st 2004



Cosmic Rays and Fundamental Physics

LUIS ANCHORDOQUI

Department of Physics

Northeastern University

Boston

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TESTBED OF NEW PHYSICS

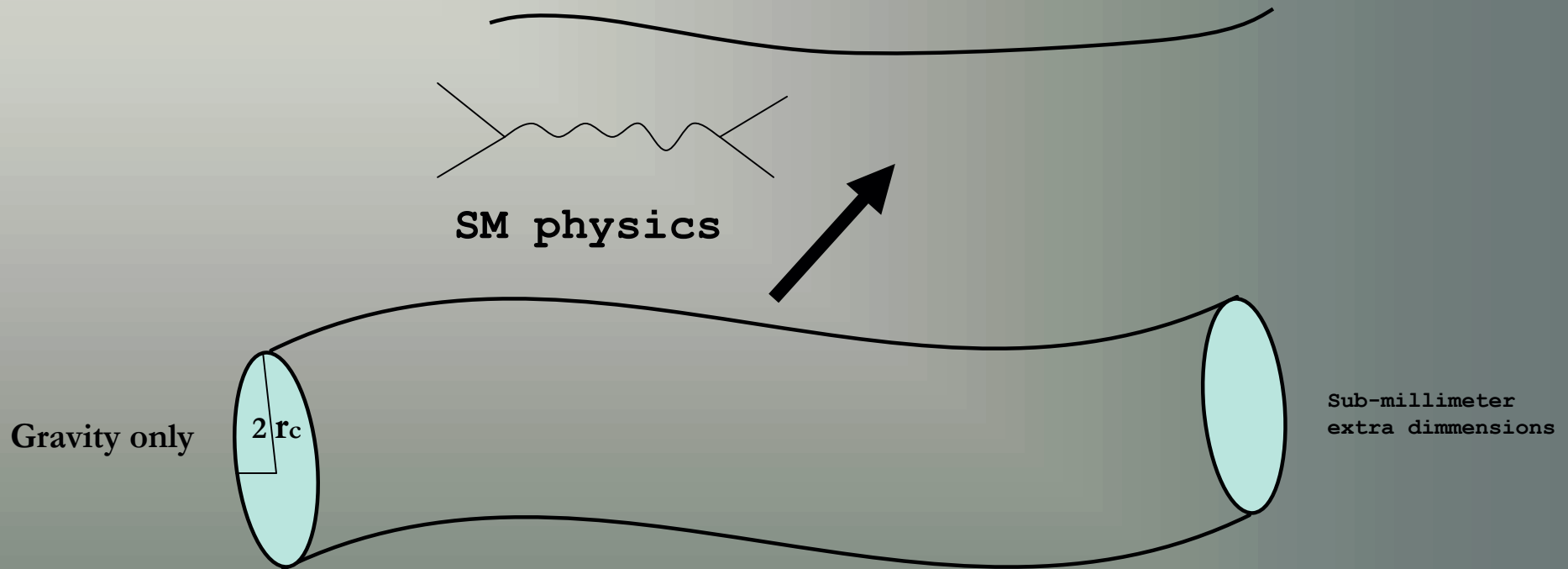




- Neutrino interactions as probes of TeV-scale gravity
- Neutrinos as messengers of high energy astrophysical processes
- Neutron β -decay as a test of local Lorentz invariance
- Probing SUSY with cosmic rays
 - » observing neutralinos from space-based detectors
 - » gluino air showers as a signal of split SUSY



Hypothesis: Universe has $D= 4 + n$ dimensions



- SM lives in 4 dimensions
- Gravity spills into internal dimensions

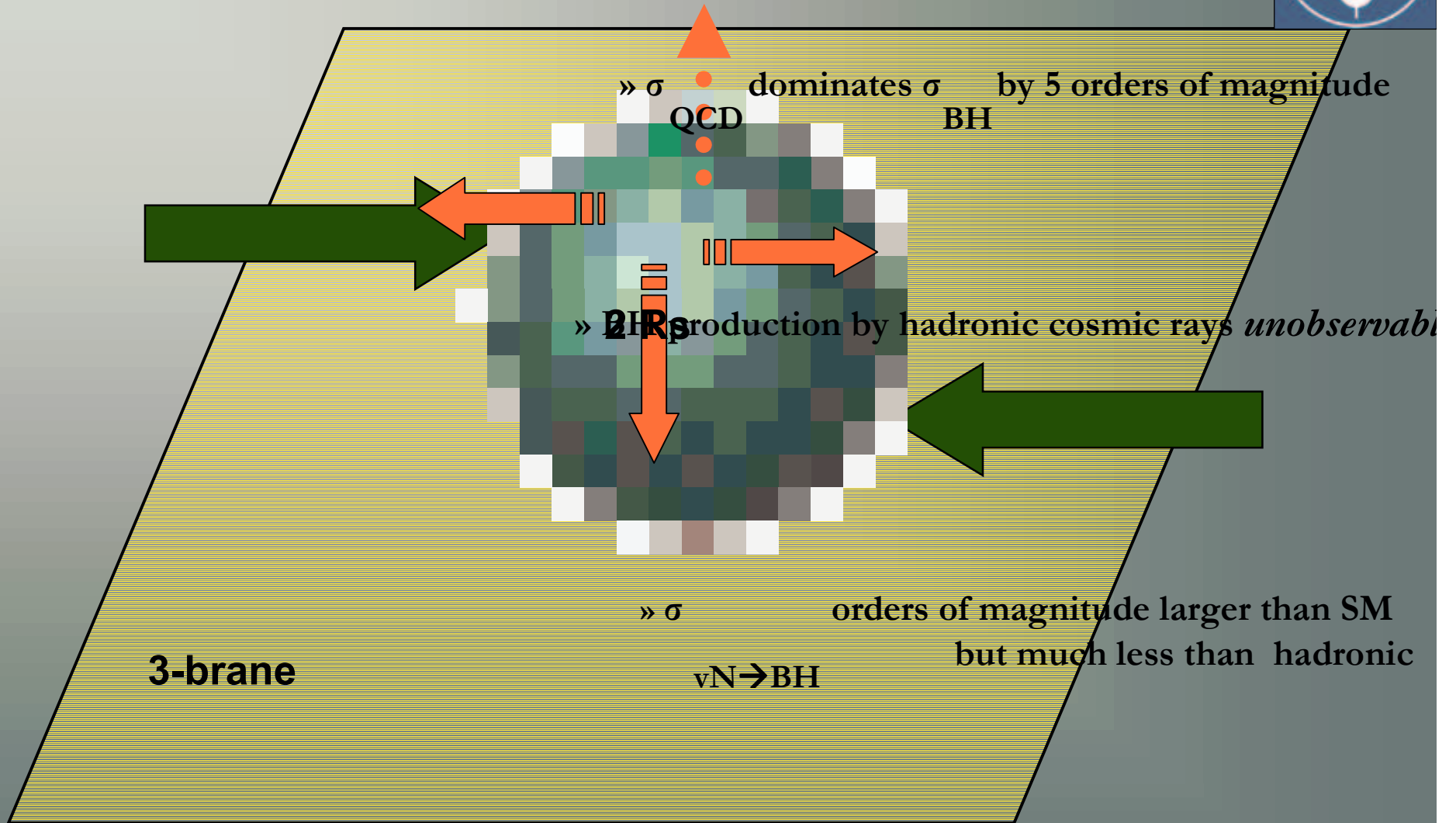


Gravitational Strength: $8\pi M_D^{2+n} r_c^n \approx (10^{19} \text{ GeV})^2$

Gravity is weak because:

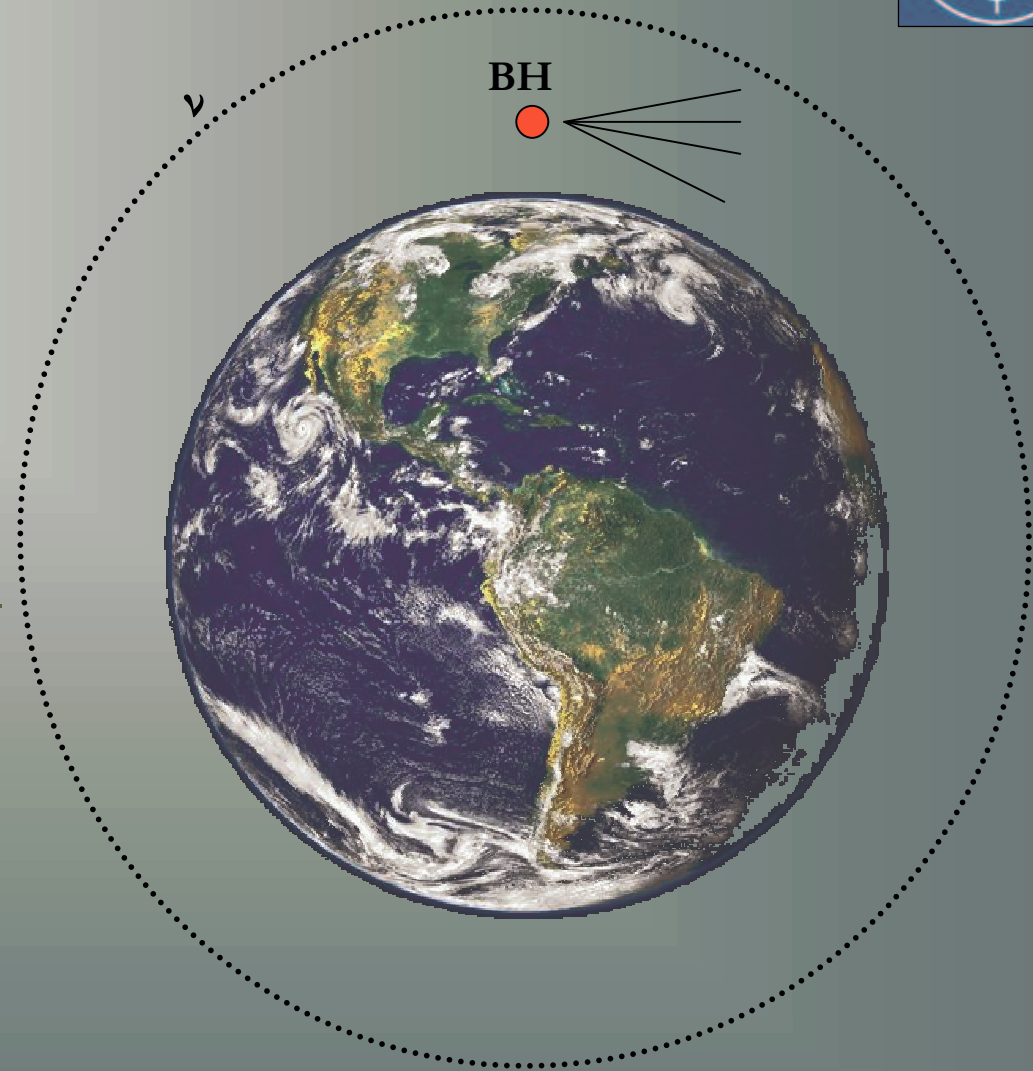
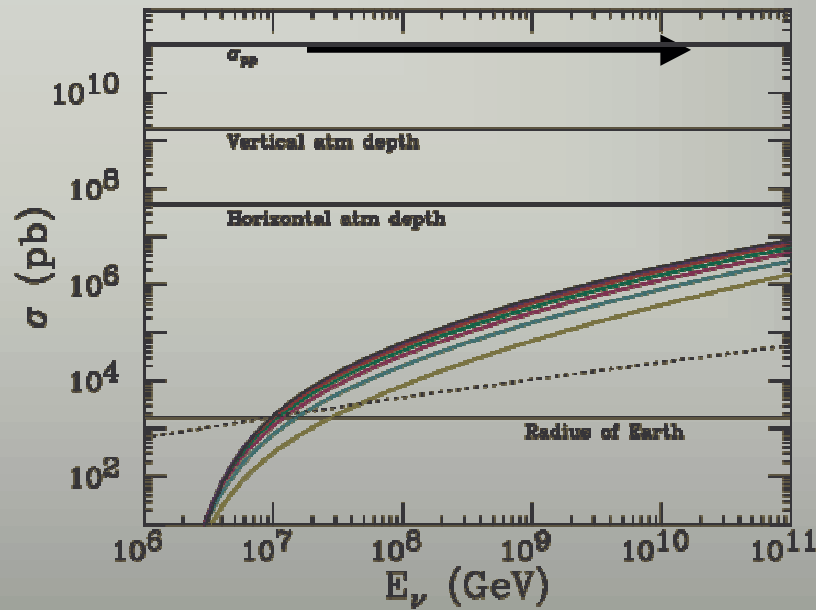
• $D = 4$ and $M_D \approx 10^{19} \text{ GeV}$

• $D > 4$ and $M_D \approx 10^3 \text{ GeV}$ but r_c large in Planck units



Dimopoulos – Landsberg PRL 87 (2001) 161602

Guidings- Thomas PRD 65 (2002) 056010



- Nearly horizontal showers are especially interesting
 - Baryonic background largely filtered by atmosphere



For Quasi-Horizontal neutrinos

- $\sigma_{\text{BH}} \Rightarrow$ increases event rates
- $\Phi > \Phi_{\text{cosmogenic}} \Rightarrow$ increases event rates

For Earth-skimmers

- $\Phi > \Phi_{\text{cosmogenic}} \Rightarrow$ increases event rates
- $\sigma_{\text{BH}} \Rightarrow$ rate suppressed

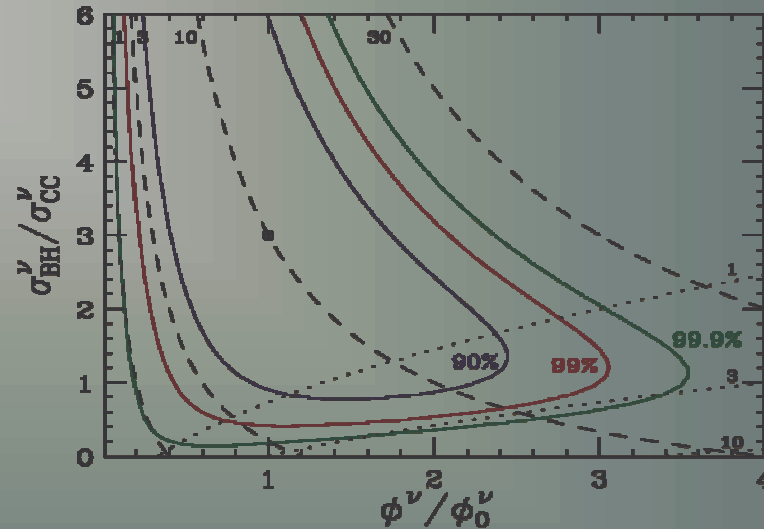
$$N_{\text{QH}} \propto \phi^\nu (\sigma_{\text{CC}}^\nu + \sigma_{\text{BH}})$$

$$N_{\text{ES}} \propto \phi^\nu \frac{\sigma_{\text{CC}}^{\nu^2}}{(\sigma_{\text{CC}}^\nu + \sigma_{\text{BH}})^2}$$

PAO 5 years

QH showers dashed

ES dotted





NON OBSERVATION OF DEEPLY PENETRATING SHOWERS



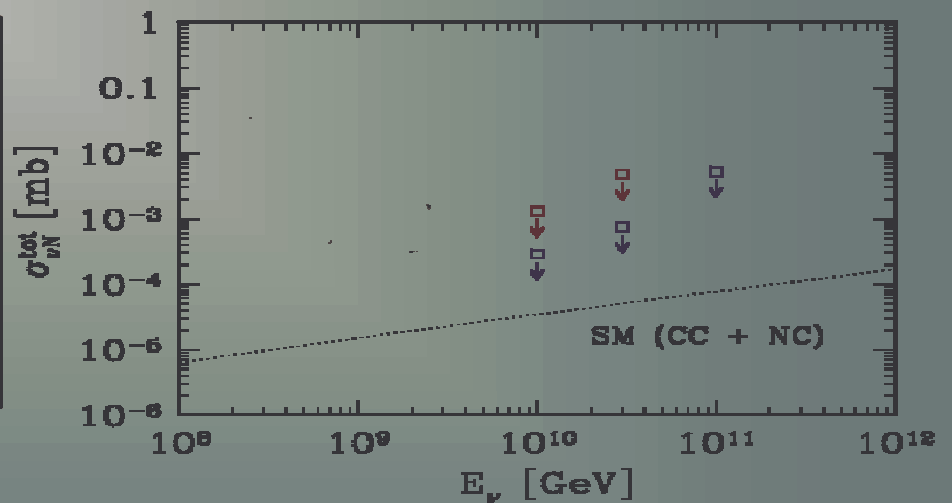
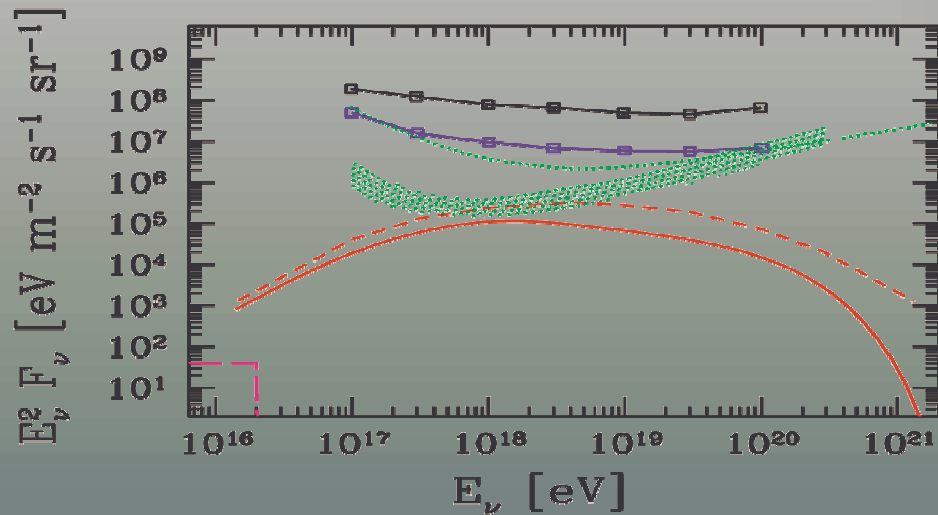
ASSUMPTION ON NEUTRINO FLUX



UHECR interactions with CMB



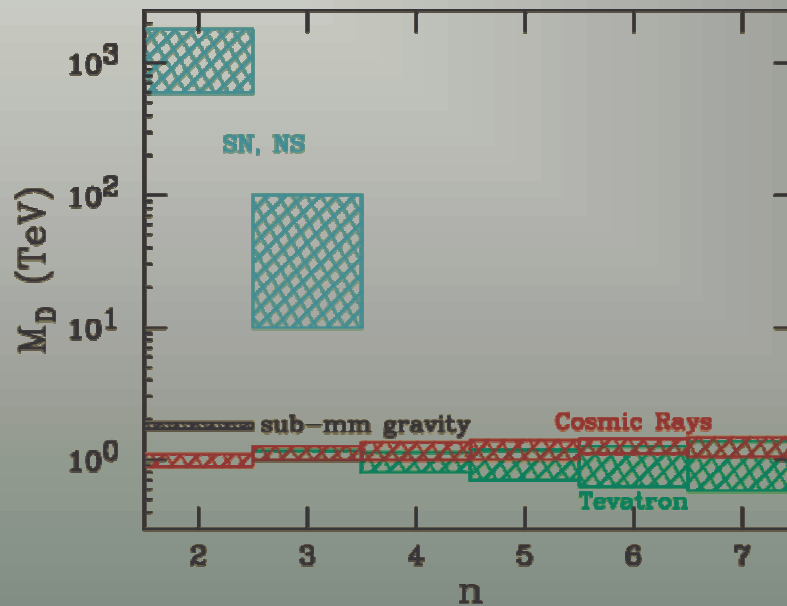
cosmogenic neutrinos



LAA-Fodor-Katz-Ringwald-Tu (*work in progress*)



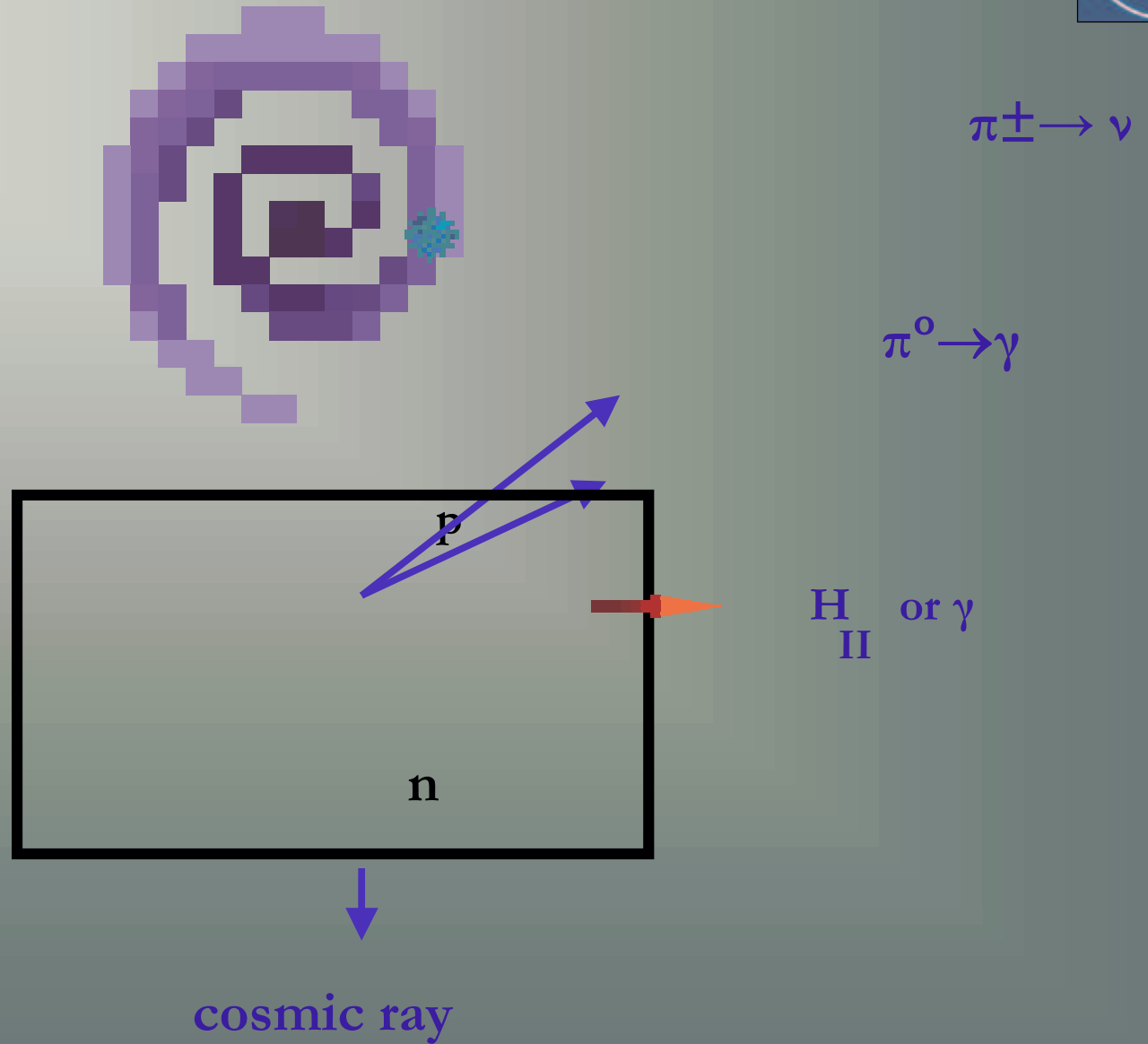
$$\sigma_{\text{BH}} \sim M_{\text{D}}^{-2}$$

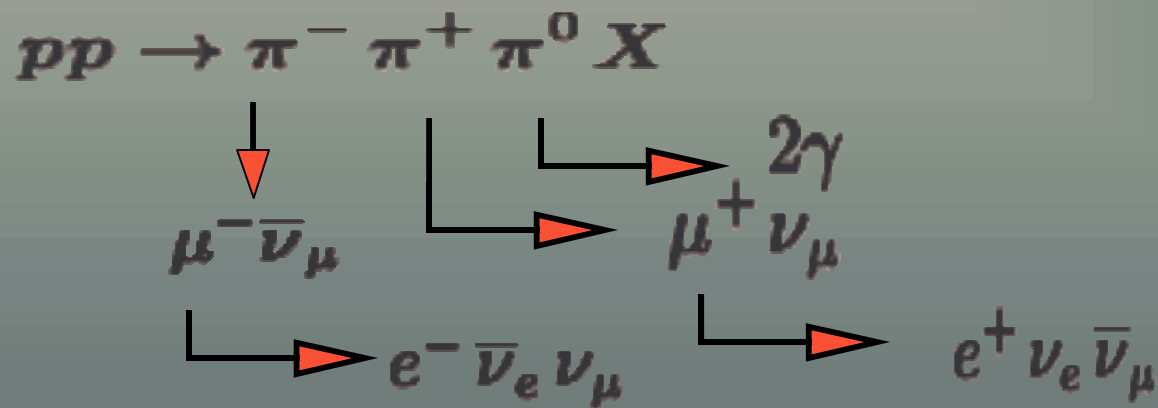
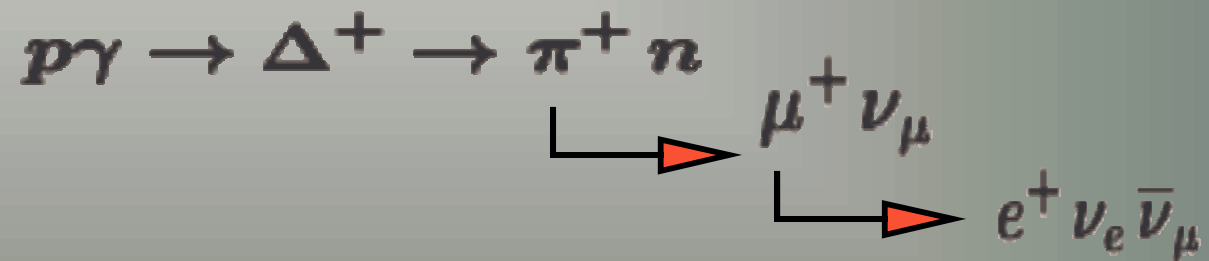


Hannestad-Raffelt PRL 88 (2002) 071301

LAA-Feng-Goldberg-Shapere PRD 68 (2003) 104025

D0 Collaboration PRL 86 (2001) 1156







Isotopically neutral mix of pions create on decay



neutrino population in the ratio

$$N_{\nu_\mu} = N_{\bar{\nu}_\mu} = 2N_{\nu_e} = 2N_{\bar{\nu}_e}$$

Neutrino oscillations

➤ Maximal $\nu_\mu \leftrightarrow \nu_\tau$ mixing

➤ $|\langle \nu_e | \nu_3 \rangle|^2 \ll 1$

➤ $\nu_3 \simeq (\nu_\mu + \nu_\tau) / \sqrt{2}$

$$N_{\bar{\nu}_e}^{\text{Earth}} = \frac{1}{6} N_{\nu+\bar{\nu}}^{\text{total}}$$



Photopion production \rightarrow isotopically asymmetric process

$$N_{\nu_{\mu}} = N_{\bar{\nu}_{\mu}} = N_{\nu_e} \gg N_{\bar{\nu}_e}$$

Neutrino oscillations

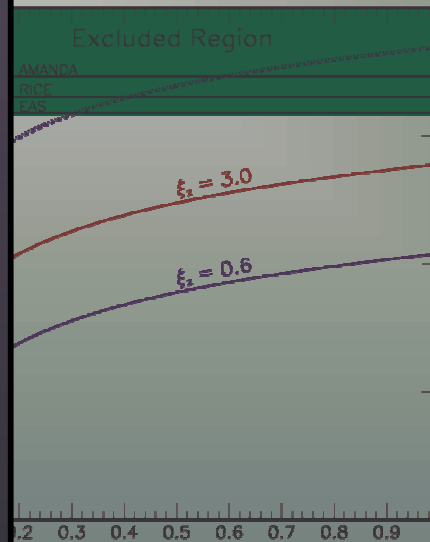
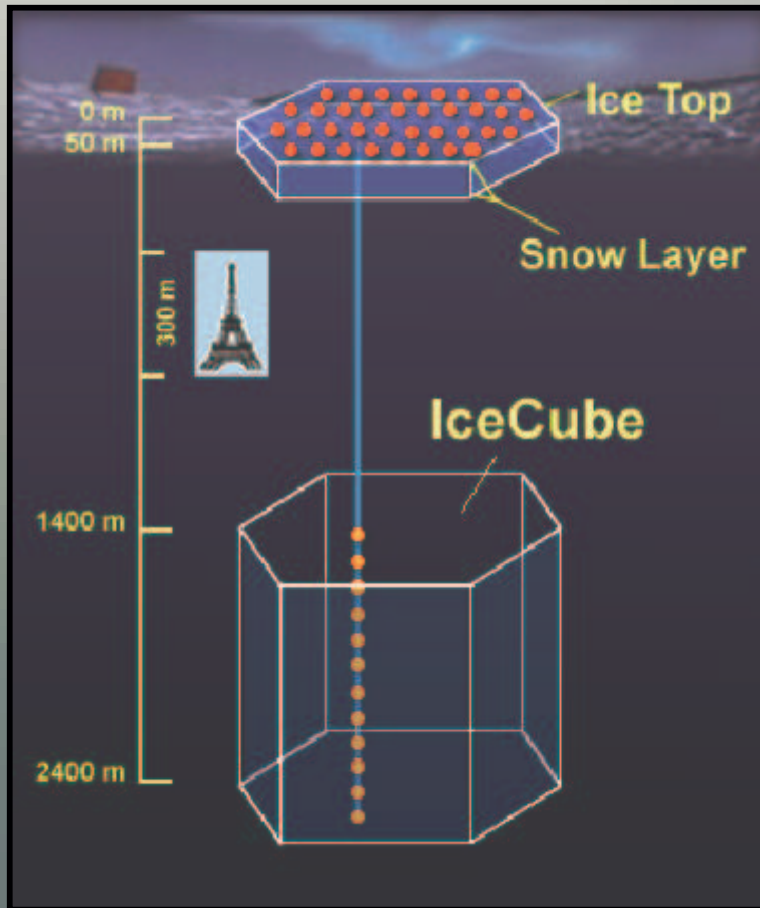
$$N_{\bar{\nu}_e}^{\text{Earth}} = N_{\bar{\nu}_{\mu}} P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e) = \frac{1}{3} \sin^2 \theta_{\odot} \cos^2 \theta_{\odot} N_{\nu+\bar{\nu}}^{\text{total}}$$

$$\text{SNO} : \theta_{\odot} \simeq 32.5^{\circ}$$

$$N_{\bar{\nu}_e}^{\text{Earth}} = \frac{1}{15} N_{\nu+\bar{\nu}}^{\text{total}}$$



Resonant scattering



Signal

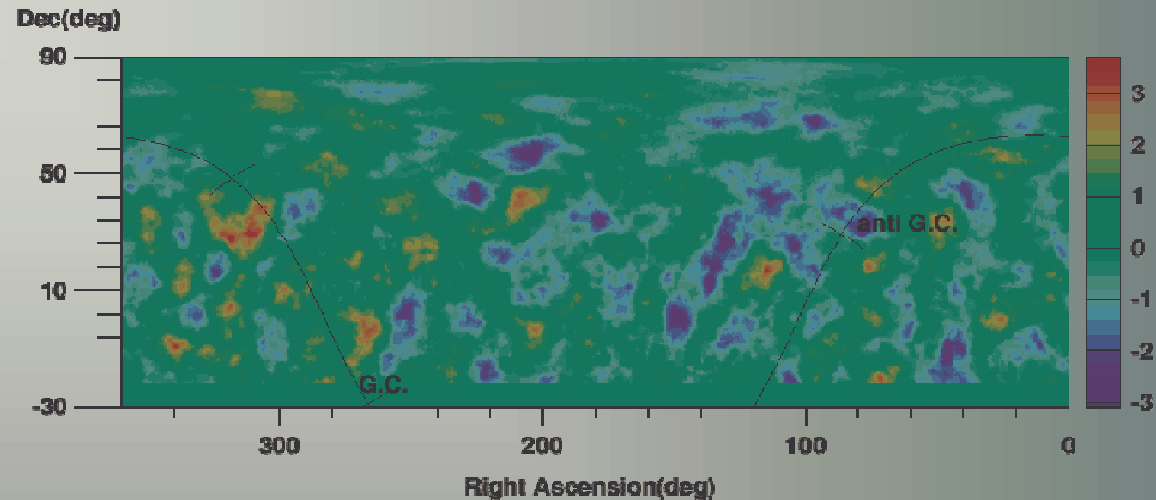
$$dN/dt = 3.2 \text{ yr}^{-1}$$

Background

$$dN/dt = 0.8 \text{ yr}^{-1}$$



Deviation of event density in equatorial coordinates as seen by AGASA



AGASA Collaboration, ICRC2001

AGASA

Largest deviation observed at coordinates $\alpha \approx 313^\circ$ $\sigma \approx 32^\circ$

4 σ deviation \Rightarrow 4 % of total flux

Fly's Eye [[astro-ph/9806096](https://arxiv.org/abs/astro-ph/9806096)]

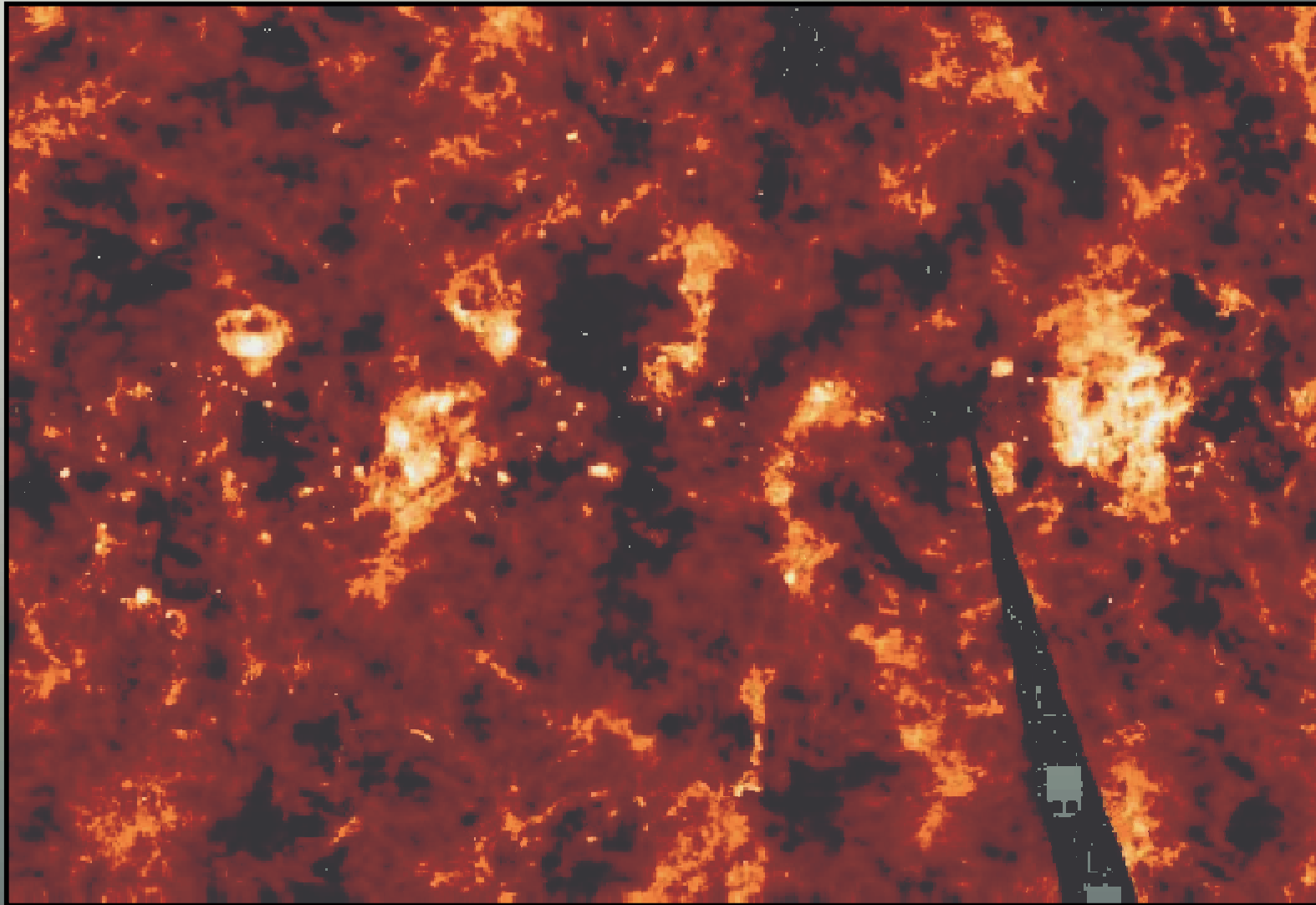
Galactic Plane enhancement \Rightarrow 3.2 σ \Rightarrow same *E* region

SUGAR [[astro-ph/0009039](https://arxiv.org/abs/astro-ph/0009039)]

Galactic Center anisotropy \Rightarrow 3 σ \Rightarrow same *E* region

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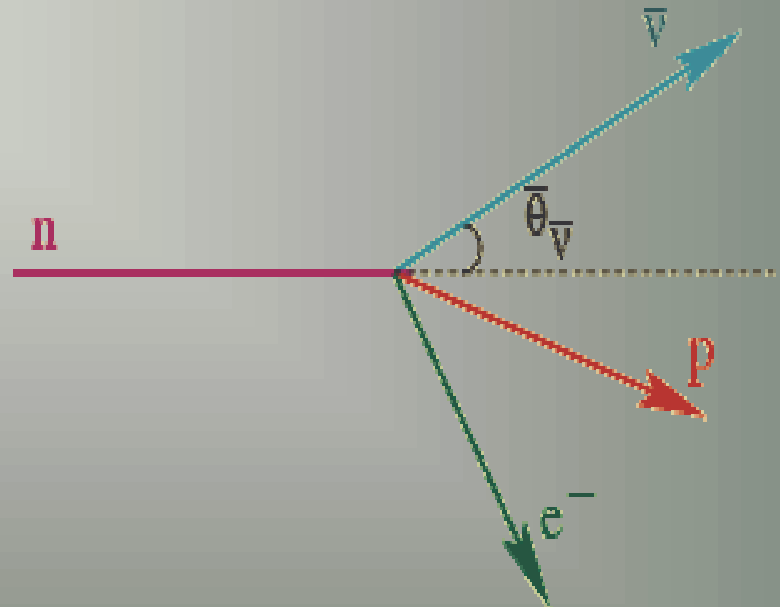
THE CYGNUS ACCELERATOR



http://antwrp.gsfc.nasa.gov/apod/ap_970424.html

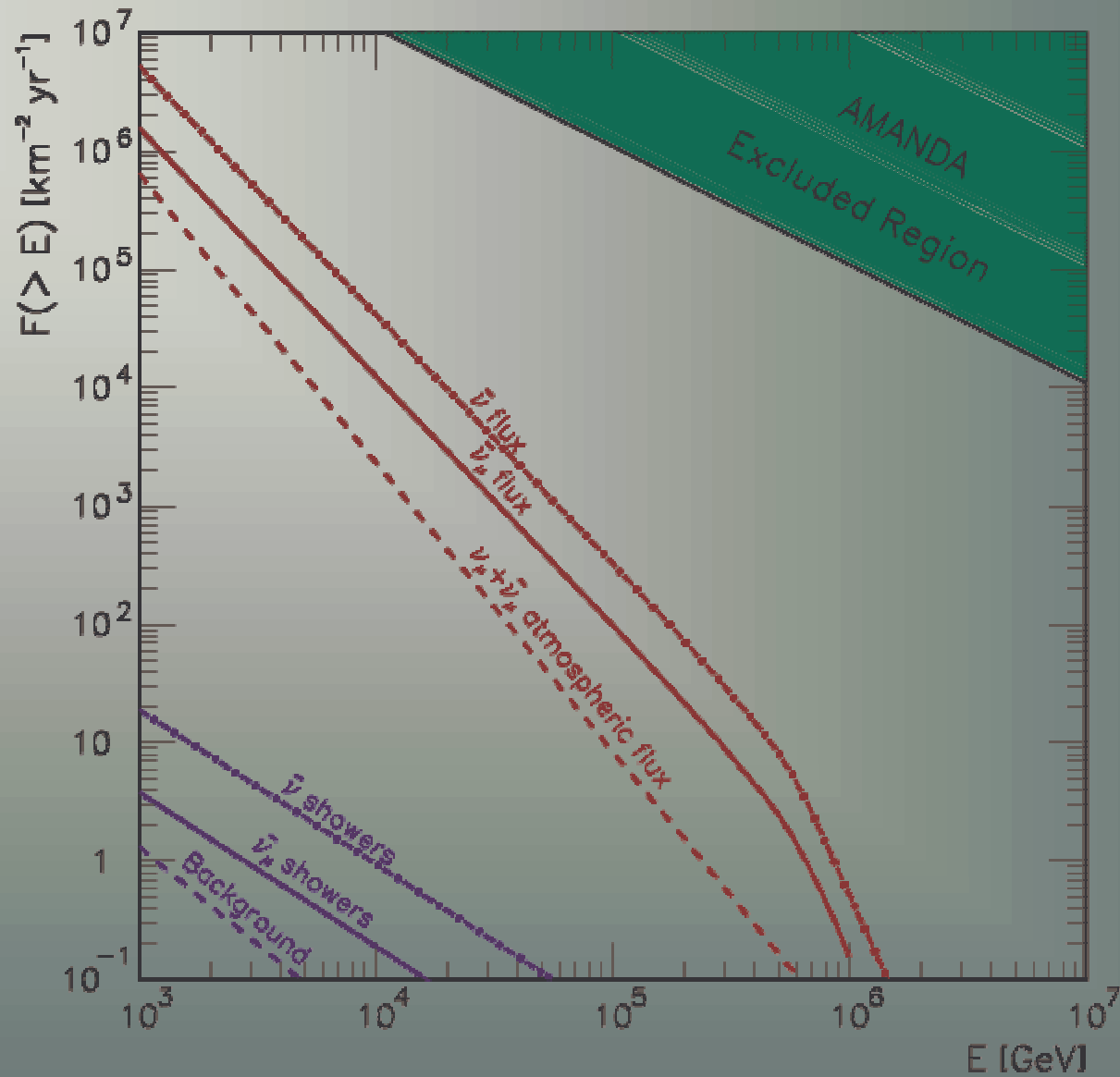


$n \text{ DK} \rightarrow \bar{\nu}$



- **Probe of Local Lorentz Invariance**
 - ☛ Shape of lower cutoff in observed anisotropy (if taken as neutrons) \Rightarrow validity of time dilation formula up to $\gamma \sim 10^9$
 - ☛ $\bar{\nu}$ -flux “smoking ice” of n -hypothesis

THE CYGNUS SIGNAL





SM particles have partners with $\Delta J = 1/2$

broken symmetry $\rightarrow \Lambda_{\text{SUSY}}$

MSSM

R – parity  stable dark matter candidate



the neutralino

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SUSY IN THE SKY



with rising energy



(2003) 132

LAA-Goldberg-Nath PRD 70 (2004) 025014



DIVINE INTERVENTION

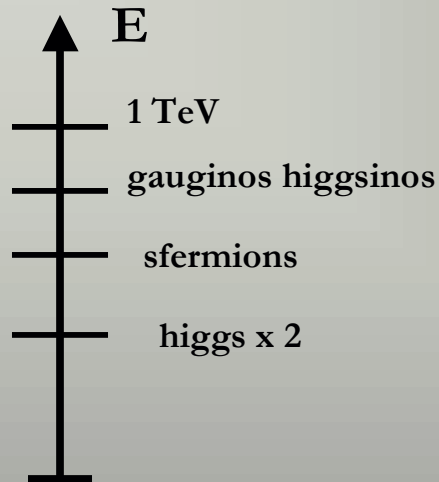


STATISTICS





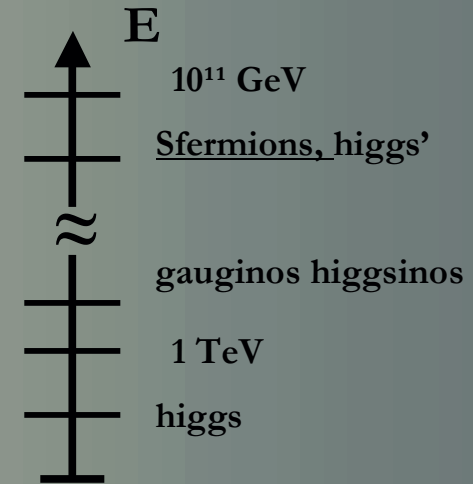
SPLIT SUSY



MSSM

The mass of the higgs is natural

Unified running



SPLIT SUSY

The mass of the higgs is fine tuned

Unification works (better)

Arkani Hamed – Dimopoulos hep-th/0405159

Giudice- Romanino hep-ph/ 0406088



➤ Because of the large mass of the sfermions \rightarrow metastable gluinos

$$\tau_{\tilde{g}} \propto m_{\tilde{g}}^4$$

➤ Very strong limits on heavy isotope abundance



Upper limit on Λ_{SUSY}

➤ Observation of cosmological gluinos



Lower limit on Λ_{SUSY}

$$10^{11} \text{ GeV} < \Lambda_{\text{SUSY}} < 10^{13} \text{ GeV}$$



SUMMARY

- ◆ Future Cosmic Ray data will not only provide clues to cosmic ray origin, but could enhance our understanding of fundamental particle physics
- ◆ The puzzle of ultrahigh energy cosmic rays may even have something to say about issues as fundamental as local Lorentz invariance
- ◆ Contrasting the observed quasi-horizontal neutrino flux with the expected neutrino flux can help to improve existing limits on the fundamental Planck scale
- ◆ An optimist might even imagine the discovery of microscopic black holes, the telltale signature of the universe's unseen dimensions
- ◆ Λ_{SUSY} written in the sky?
- ◆ We are entering this new High Energy Physics era with PAO + Ice Cube

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THANKS

&

QUESTIONS

