Current status and future prospects of the singlet-doublet dark matter model with *CP* violation

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We discuss the singlet-doublet fermion dark matter model with CP violation. In this model, the CP violation generates a pseudoscalar interaction of dark matter with the standard model Higgs boson. Thanks to the pseudoscalar interaction, the model can evade the strong constraint from the dark matter direct detection experiments while keeping the success of the thermal relic scenario. The CP violation also predicts signals in dark matter indirect detection experiments and electric dipole moments (EDM) that can be as large as the current upper bound. We investigate the constraints and prospects of the direct detection experiments, the indirect detection experiment, and the electron EDM. We also investigate the stability of the Higgs potential. Combining these observables, we find that heavy dark matter is disfavored. We also find it is possible to probe the Higgs funnel region by the combination of the direct detection experiments and the measurements of the electron EDM if experiments for the electron EDM reach to $\mathcal{O}(10^{-32}) \ e$ cm in future.

1. INTRODUCTION

The thermal relic scenario is an attractive scenario to explain the origin of the dark matter (DM). This scenario does not rely on the initial condition of our Universe, and it has many indications on model building of DM. See for a review Ref. [1]. In this scenario, the energy density of the DM in the Universe is determined by the annihilation of the DM particles into the standard model (SM) particles. Therefore, this scenario requires DM-SM interactions, and we can expect nonzero DM-nucleon scattering cross section. There are experiments that aim to detect the DM directly through this scattering process, such as the Xenon1T [2] and PandaX-II experiments [3]. However, signals of the DM-nucleon scattering have not been detected yet. This null result gives severe

constraints on the DM models, and it becomes important to investigate models which can avoid such constraints.

The simplest model which can avoid the constraints from the direct detection experiments is an effective theory that is constructed by introducing a gauge singlet fermion DM [4–11].¹ The DM in this setup does not couple to the SM field via renormalizable interactions. The DM-Higgs interactions are realized by dimension-5 operators, $\bar{\psi}\psi H^{\dagger}H$ and $\bar{\psi}i\gamma^5\psi H^{\dagger}H$. If the coefficient of the *CP*-conserving operator $\bar{\psi}\psi H^{\dagger}H$ is negligible, then the *CP*-violating operator $\bar{\psi}i\gamma^5\psi H^{\dagger}H$ controls both of the amount of the relic abundance and the DM-nucleon scattering cross section. The *CP*-violating operator generates a pseudoscalar interaction between DM and the SM Higgs boson, $\bar{\psi}i\gamma^5\psi h$. Since the nonrelativistic DM-nucleon scattering cross section with this interaction is highly suppressed by the relative velocity between the DM and the nucleon, we can avoid the constraints from the direct detection experiments while keeping the amount of the relic abundance [5]. The interaction term is nonrenormalizable, and a cutoff scale accompanies it. The value of the cutoff scale to be required to obtain the correct relic abundance is $\mathcal{O}(1)$ TeV. This result motivates us to consider UV completions of the effective theory as physics at TeV scale.

The singlet-doublet DM model [13–15] is one of the UV completion of the above effective theory. The model contains a gauge singlet Majorana fermion and an SU(2) doublet Dirac fermion with hypercharge 1/2. This setup can be regarded as a generalization of the Bino-Higgsino system in the minimial supersymmetric standard model [16–19]. In the large Dirac mass limit, the model is reduced into the effective theory as schematically shown in Fig. 1. The model contains a CP phase, and thus the pseudoscalar interaction exists. Although this model has been widely studied [19–34], its DM phenomenology with CP violation has not been studied sufficiently. See Refs. [35, 36] for previous studies. The DM annihilation processes with the *s*-channel exchange of the Higgs boson with the pseudoscalar interaction are *s* wave, and its cross section at low temperature is not suppressed by powers of velocity. As a result, the annihilation cross section of the DM in the current Universe can be sizable enough to generate some signals in cosmic rays such as γ rays. Dwarf spheroidal galaxies (dSphs) are regarded as good targets to observe such γ -ray emissions because of the less astrophysical uncertainty. Using the Fermi-LAT data [37] that measures the γ -ray flux from dSphs, we can investigate the CP violation in the singlet-doublet DM model.

Another consequence of the CP violation is electric dipole moments (EDMs). The predicted value of the electron EDM (eEDM) by the SM is $|d_e| \leq 10^{-38} \ e \ cm$ [38, 39]. This is much smaller

¹ A similar but nonminimal approach by using a two-Higgs doublet model with Z_3 was discussed in Ref. [12].



FIG. 1: The schematic view of the UV completion from the effective theory to the singlet-doublet model.

than the current upper bound on the eEDM by the ACME experiment [40], $|d_e| \leq 1.1 \times 10^{-29} e$ cm (90% C.L.). It is known that the singlet-doublet model predicts the eEDM that can be as large as the current upper bound [13, 14, 35, 36]. Since the constraints from the direct detection experiments and the measurements of the eEDM are strong enough to probe the model, it is worth studying the correlation among the eEDM and other observables, such as the DM-nucleon scattering cross section, to see if the model provides a viable DM candidate.

The stability of the Higgs potential also gives a bound on the singlet-doublet model. Higgs couplings to fermions can give negative contributions to the beta function of the Higgs quartic coupling and make the Higgs potential unstable [41]. While the pseudoscalar interaction of the DM with the Higgs boson is essential in the singlet-doublet model to avoid the strong constraint from the direct detection experiments, it can make the Higgs potential unstable. Therefore, imposing the stability on the Higgs potential gives another constraint on the singlet-doublet model.

The rest of this paper is organized as follows. In Sec. 2, we briefly review the current status of the minimal fermionic dark matter model. We show the constraints on the *CP*-conserving and *CP*-violating DM-Higgs boson interactions. It is found that the *CP*-conserving interaction is severely constrained from the DM direct detection, and the *CP* violation implies that a UV complete model appears around $\mathcal{O}(1)$ TeV scale. In Sec. 3, we discuss the singlet-doublet fermion model. We focus on the constraint from the direct detection, indirect detection, the stability of the Higgs potential, and the eEDM. The Higgs invisible decay is also discussed for $m_{\rm DM} < m_h/2$. It is shown that the heavy dark matter is disfavored by the combination of the constraints. We also show that most of the region of the parameter space, including the Higgs funnel region, can be probed by the combination of the direct detection experiments and the measurements of the eEDM in future. In Sec. 4, we present our conclusion.

2. EFFECTIVE THEORY FOR FERMIONIC DM MODELS

In this section, we discuss a Z_2 -odd gauge singlet Majorana fermion (ψ) as a DM candidate. If we do not introduce any other new particle, the DM cannot couple to the SM particles at the renormalizable level and we need higher dimensional operators. We can write dimension-five operators with the Higgs boson.

$$\mathcal{L} = \mathcal{L}_{\rm SM} + \frac{1}{2} \bar{\psi} \left(i \gamma^{\mu} \partial_{\mu} - m_{\psi} \right) \psi + \frac{c_S}{2} \bar{\psi} \psi \left(H^{\dagger} H - \frac{v^2}{2} \right) + \frac{c_P}{2} \bar{\psi} i \gamma_5 \psi \left(H^{\dagger} H - \frac{v^2}{2} \right), \qquad (2.1)$$

where H is the SM Higgs field. There are two operators, $\bar{\psi}\psi H^{\dagger}H$ and $\bar{\psi}i\gamma_5\psi H^{\dagger}H$. The former is a CP-conserving operator, and the latter violates the CP invariance. The CP-conserving interaction has been studied in Ref. [4], and the CP-violating operator has been studied in Refs. [5–7, 9, 11].

These two operators have different properties in the DM physics. In nonrelativistic DM-nucleon scattering processes, only the CP-conserving operator is relevant. Thus, the spin-independent cross section is proportional to c_S^2 ,

$$\sigma_{\rm SI} = \frac{1}{\pi} \frac{f_N^2 c_S^2}{m_h^4} \frac{m_N^4 m_{\rm DM}^2}{(m_N + m_{\rm DM})^2} \simeq \left(\frac{c_S}{\text{TeV}^{-1}}\right)^2 3.22 \times 10^{-44} \text{ cm}^2, \qquad (2.2)$$

where m_N is the nucleon mass. Here we have used the following numbers.

$$m_N = 0.938 \text{ GeV},$$
 (2.3)

$$f_N = \frac{2}{9} + \frac{7}{9} \sum_q f_q, \qquad (2.4)$$

$$f_u = 0.0110, \quad f_d = 0.0273, \quad f_s = 0.0447.$$
 (2.5)

where f_q values are taken from micrOMEGAs [42].

On the other hand, both operators contribute to the DM annihilation processes, which are important to determine the dark matter relic density. The annihilation cross section is the p-wave (s-wave) process with the CP-conserving (CP-violating) operator.

Figure 2 shows the parameters which explain the DM relic density $\Omega h^2 = 0.1198 \pm 0.0015$ [43] in the c_S - c_P plane. We find the result is almost independent from c_S . This is because the *CP*conserving and the *CP*-violating operators contribute to the DM annihilation processes by p wave and s wave, respectively, and thus the *CP*-violating operator plays the dominant role in the determination of the DM thermal relic abundance. We also find $c_P \sim 0.2$ TeV⁻¹ except for the Higgs funnel region. This c_P value implies that UV completions of the effective Lagrangian contain one or more new particles around $\mathcal{O}(1)$ TeV.



FIG. 2: The fermion dark matter with $m_{\rm DM} = 55$ (upper left), 100 (upper right), 300 (lower left), and 1000 GeV (lower right). The relic abundance $\Omega h^2 = 0.1198$ on the red lines. The right region of the blue lines is excluded by the XENON1T experiment [2].

Figure 3 shows the spin-independent cross section for the weakly interacting massive particle (WIMP)-nucleon scattering process as a function of the DM mass. Here we fix a parameter so as to obtain the measured value of the DM energy density. If the DM candidate interacts with the SM particles only through the *CP*-conserving operator, namely $c_P = 0$, then the constraint from the DM direct detection experiments excludes the large part of the parameter space except for the Higgs funnel region. On the other hand, if we have the *CP*-violating interaction ($c_P \neq 0$), then it is easy to avoid the constraints from the current DM direct detection experiments as can be seen from the figure. Therefore, we need nonzero c_P except for the Higgs funnel region.

The DM annihilation processes with c_P are s wave except for the Higgs funnel region, and thus the annihilation cross section at low temperature is not suppressed by powers of velocity. This leads $\langle \sigma v \rangle \sim \mathcal{O}(10^{-26}) \text{ cm}^3/s$ in the current Universe. As a result, we can expect the DM annihilation signals in the current Universe. We focus on γ -ray emissions by the DM annihilation



FIG. 3: The spin-independent cross section for the DM-nucleon scattering. The current upper bound on $\sigma_{\rm SI}$ from the XENON1T experiment [2] is shown by the black solid line. For $m_{\rm SM} > 1$ TeV, we extrapolate the result given in Ref. [2], and show it by the black dashed line. We also show the future prospects for XENONnT experiment [44] and LZ experiment [45] by the red solid line. The other lines are the prediction of the effective theory given in Eq. (2.1).

from dSphs, which are regarded as good targets to observe such γ -ray emissions because of the less astrophysical uncertainty. The γ -ray flux from DM annihilation is given by [37]

$$\phi(\Delta\Omega, E_{\min}, E_{\max}) = \frac{1}{4\pi} \frac{\langle \sigma v \rangle}{2m_{\rm DM}^2} \int_{E_{\min}}^{E_{\max}} \frac{dN_{\gamma}}{dE_{\gamma}} dE_{\gamma} \times J, \qquad (2.6)$$

where J is the J-factor given by

$$J = \int_{E_{\min}} \int_{\Delta\omega} \int_{l.o.s} \rho_{\rm DM}^2(\vec{r}(\ell)) d\ell d\Omega.$$
(2.7)

Each dSph has different J-factor, and some of their J-factors are already measured. The variables other than the J-factor in Eq. (2.6) are calculable in particle physics. We investigate the Fermi-LAT gamma-ray data [37], and use the likelihood functions.² We use the 19 dSphs whose J-factors were measured and are listed in the sixth column in Table 1 in Ref. [37]. With these J-factors,

 $^{^2}$ The likelihood functions for each dSph are given at http://www-glast.stanford.edu/pub_data/1203/.



FIG. 4: The constraint on the effective theory from the measurements of the γ -ray flux from the dSphs by the Fermi-LAT experiment. In the upper panel, we use the 19 dSphs whose J-factors were measured kinematically.

we use micrOMEGAs [42] to calculate the γ -ray flux from the DM annihilation. Figure 4 shows the model prediction on $\langle \sigma v \rangle$ for $c_S = 0$ with the constraints from the indirect detection. We find the upper bound on $\langle \sigma v \rangle$ and the regions for $m_{\rm DM} < 50$ GeV and for 62 GeV $< m_{\rm DM} < 78$ GeV are excluded.

There are two important points from Figs. 2–4. First, we need the CP-violating operator except for the funnel region. Second, the cutoff scale of this effective theory is around O(1) TeV. This means a UV completion of this model requires one or more new particles at the TeV scale in addition to the DM particle. In the next section, we focus on the singlet-doublet DM model [13–15] as a UV completion of the effective theory discussed in this section.

3. THE SINGLET-DOUBLET MODEL

In this section, we discuss the singlet-doublet model [13–15] as a UV completion of the effective theory which we have discussed in the previous section. After reviewing the setup of the singletdoublet model briefly, we discuss the constraint on the DM annihilation cross section from the Fermi-LAT experiment [37]. Since the DM annihilation process is s wave due to the CP-violating operator, the constraint is expected to be stronger than the case without CP violation in the dark sector [27]. The *CP*-violating operator also generates EDMs. We focus on the electron EDM and discuss its complementary role to the DM direct detection searches. We also investigate the stability of the Higgs potential. In the singlet-doublet model, the Higgs boson couples to the new fermion fields in the dark sector. The couplings give the negative contribution to the beta function of the Higgs quartic coupling. As a result, the Higgs potential becomes unstable compared to the Higgs potential in the SM.

Before starting to discuss the singlet-doublet model, we make brief comments on other UV completions. A model discussed in Refs. [46–48], a gauge singlet fermion and a gauge singlet CP-odd scalar are introduced. The singlet fermion is a DM candidate, and the CP-odd scalar is a mediator particle. In the heavy CP-odd scalar mass limit, the model can be described by the effective theory discussed in Sec. 2. Another UV completion where the Higgs sector is also extended into the two-Higgs doublet model was proposed in Ref. [49] and has been widely studied [50–59]. For other models, see Ref. [24].

3.1. The model

We introduce a gauge singlet Majorana fermion (ω) and an $\mathrm{SU}(2)_L$ doublet Dirac fermion with hypercharge Y = 1/2 that consists of a left-handed Weyl fermion $(\eta = (\eta^+, \eta^0)^T)$ and a righthanded Weyl fermion $(\xi^{\dagger} = ((\xi^-)^{\dagger}, \xi^{0\dagger})^T)$; see Table I. We impose a Z_2 symmetry on the model. Under the Z_2 symmetry, all the SM particles are even, and all the fermion fields we introduced in the above are odd.

	$\mathrm{SU}(3)_c$	$\mathrm{SU}(2)_L$	$\mathrm{U}(1)_Y$	Z_2
ω	1	1	0	-1
η	1	2	1/2	-1
ξ^{\dagger}	1	2	1/2	-1

TABLE I: New particles in the Weyl notation.

The mass and Yukawa interaction terms for the Z_2 -odd particles are given by

$$\mathcal{L}_{int.} = -\frac{M_1}{2}\omega\omega - M_2\xi\eta - y\omega H^{\dagger}\eta - y'\xi H\omega + (h.c.), \qquad (3.1)$$

where M_1 , M_2 , y, and y' are complex parameters. Three of the complex phases are unphysical because they can be absorbed by rotating the Z_2 -odd fermion fields. The physical phase (CP phase) is given by

$$\phi = \operatorname{Arg}(M_1^* M_2^* y y'). \tag{3.2}$$

We introduce r for later convenience,

$$r = \frac{|y|}{|y'|}.\tag{3.3}$$

We have five free parameters, $(|M_1|, |M_2|, \sqrt{|y|^2 + |y'|^2}, r, \phi)$. We determine $\sqrt{|y|^2 + |y'|^2}$ so as to realize the measured value of the dark matter energy density, $\Omega h^2 = 0.1198 \pm 0.0015$ [43]. We use **micrOMEGAs** [60] to calculate the dark matter relic abundance. After the Higgs field develops the vacuum expectation value, the singlet field and the doublet field are mixed. As a result, there are three Majorana fermions $(\chi^0_{1,2,3})$ and a pair of charged fermions (χ^{\pm}) in the Z₂-odd sector. The lightest Majorana fermion is the DM candidate in this setup.

3.2. Relation to the effective theory

In order to see the relation of this model to the effective theory, we discuss the large M_2 regime. If $M_2 \gg v, M_1$, then we can integrate out the doublet fields, and the model is reduced into the effective theory. Up to dimension-6 operators, we find the following terms in the effective Lagrangian.

$$\begin{aligned} \mathcal{L} = \mathcal{L}_{\rm SM} + \frac{1}{2} \Psi_s \left(i \gamma^{\mu} \partial_{\mu} - |M_1| \right) \Psi_s \\ + \operatorname{Re} \left(\frac{|yy'|e^{i\phi}}{|M_2|} \right) H^{\dagger} H \bar{\Psi}_s \Psi_s - \operatorname{Im} \left(\frac{|yy'|e^{i\phi}}{|M_2|} \right) H^{\dagger} H \bar{\Psi}_s i \gamma^5 \Psi_s \\ - \frac{|y|^2 - |y'|^2}{4|M_2|^2} \left((H^{\dagger} i \overleftrightarrow{D}_{\mu} H) \bar{\Psi}_s \gamma^5 \gamma^{\mu} \Psi_s \right) + \frac{|y|^2 + |y'|^2}{2|M_2|^2} \left(H^{\dagger} H \bar{\Psi}_s \gamma^{\mu} i \partial_{\mu} \Psi_s \right) \end{aligned}$$
(3.4)
$$= \mathcal{L}_{SM} + \frac{1}{2} \Psi_s \left(i \gamma^{\mu} \partial_{\mu} - |M_1| \right) \Psi_s \\ + \left(\operatorname{Re} \left(\frac{|yy'|e^{i\phi}}{|M_2|} \right) + M_1 \frac{|y|^2 + |y'|^2}{2M_2^2} \right) H^{\dagger} H \bar{\Psi}_s \Psi_s - \operatorname{Im} \left(\frac{|yy'|e^{i\phi}}{|M_2|} \right) H^{\dagger} H \bar{\Psi}_s i \gamma^5 \Psi_s \\ - \frac{|y|^2 - |y'|^2}{4|M_2|^2} (H^{\dagger} i \overleftrightarrow{D}_{\mu} H) \bar{\Psi}_s \gamma^5 \gamma^{\mu} \Psi_s, \end{aligned}$$
(3.5)

where

$$\Psi_s = \begin{pmatrix} \omega \\ \omega^{\dagger} \end{pmatrix}, \tag{3.6}$$

$$A \overleftrightarrow{}_{\mu} B = A \partial_{\mu} B - (\partial_{\mu} A) B.$$
(3.7)

Here we have used the equation of motion for Ψ_s . Comparing Eq. (3.5) with Eq. (2.1), we find that M_2 plays the role of the cutoff scale in the effective theory.

As can be seen from Eq. (3.5), there is a region in the parameter space where the scalar coupling is highly suppressed or even vanishes while keeping the pseudoscalar coupling. This accidental cancellation of the scalar coupling is known as the blind spot [22, 23, 25, 28, 35]. Therefore the scalar coupling can be parametrically suppressed in this model.

The last term in Eq. (3.5) is absent in Eq. (2.1) because we did not consider the effects of dimension-6 operators in the previous section. Since it generates the DM-Z coupling, it affects the DM annihilation cross section and also contributes to the spin-dependent DM-nucleon scattering process. If |y| = |y'|, namely r = 1, this term vanishes because of the symmetry enhancement in the dark sector. If r = 1, the dark sector is symmetric under the exchange of η with ξ , and consequently the DM-Z coupling vanishes. The beta function of r at the one-loop level is given by

$$\mu \frac{d}{d\mu} \frac{y}{y'} = -\frac{1}{(4\pi)^2} \frac{3y}{2y'} \left(|y|^2 - |y'|^2 \right).$$
(3.8)

It is clear that r = 1 is a fixed point because the beta function is 0. This is another way to see the symmetry enhancement at r = 1. In the following numerical analysis, we take r = 1 for simplicity.

3.3. Electron EDM

EDMs, d_f , are sensitive to CP violation and are defined through

$$\mathcal{H}_{eff} = i \frac{d_f}{2} \bar{\psi}_f \sigma_{\mu\nu} \gamma_5 \psi_f F^{\mu\nu}. \tag{3.9}$$

As we have discussed in Sec. 2, CP violation is important to avoid the constraints from the DM direct detection experiments, and thus the EDMs are naturally expected in this model [13, 14, 35, 36]. In particular, we focus on the eEDM, d_e , that is severely constrained by experiments.

The current bound on the eEDM is given by the ACME experiment [40], $|d_e| \leq 1.1 \times 10^{-29} e$ cm (90% C.L.). There are some prospects for eEDM [61, 62], and we can expect that the eEDM is detectable in future if $|d_e| \gtrsim \mathcal{O}(10^{-30}) e$ cm. We use $|d_e| = 10^{-30} e$ cm as the prospect of eEDM in the following analysis.

The eEDM in this model is given by [35, 36]

$$\frac{d_e}{e} = -\frac{2\alpha}{(4\pi)^3 s_W^2} \sqrt{2} G_F m_{\chi^{\pm}} m_e \sum_{j=1}^3 \operatorname{Im}(V_{2j} V_{3j}) m_{\chi_j^0} \mathcal{I}_j, \qquad (3.10)$$

where $m_{\chi^{\pm}}$ is the mass of the charged Z₂-odd fermion, which is the same as $|M_2|$ at the tree level, V_{ij} is a 3 by 3 matrix to connect the gauge eigenstates and the mass eigenstates of the Z₂-odd fermions,

$$\begin{pmatrix} \omega \\ \eta^0 \\ \xi^0 \end{pmatrix} = \begin{pmatrix} V_{11} & V_{12} & V_{13} \\ V_{21} & V_{22} & V_{23} \\ V_{31} & V_{32} & V_{33} \end{pmatrix} \begin{pmatrix} \chi^0_1 \\ \chi^0_2 \\ \chi^0_3 \end{pmatrix},$$
(3.11)

and

$$\mathcal{I}_{j} = \int_{0}^{1} dz \frac{1-z}{m_{\chi^{\pm}}^{2}(1-z) + m_{\chi^{0}_{j}}^{2}z - m_{W}^{2}z(1-z)} \ln \frac{m_{\chi^{\pm}}^{2}(1-z) + m_{\chi^{0}_{j}}^{2}z}{m_{W}^{2}z(1-z)}.$$
(3.12)

For $|M_2| \gg |M_1|, v$,

$$\frac{d_e}{e} \simeq -\frac{1}{2} \frac{1}{(4\pi)^4} \left(\frac{e}{s_W}\right)^2 m_e \frac{|M_1 M_2 y y'| \sin \phi}{|M_2|^4} \ln \frac{|M_2|^2}{m_W^2}.$$
(3.13)

It is clear that the eEDM is proportional to $\text{Im}(M_1^*M_2^*yy')$. For $|M_2| \gg |M_1|, v$, we find $m_{\text{DM}} \simeq M_1$, and thus the EDMs are proportional to the dark matter mass.

3.4. Stability of the Higgs potential

Before studying the DM physics, we discuss the stability of the Higgs potential.³ It is worse than the SM case because of the Higgs couplings to the fermion fields in the dark sector [63]. The beta function of the Higgs quartic coupling at the one-loop level is given by

$$(4\pi)^{2} \mu \frac{d\lambda}{d\mu} = +24\lambda^{2} + \frac{3}{8}g_{1}^{4} + \frac{3}{4}g_{1}^{2}g_{2}^{2} + \frac{9}{8}g_{2}^{4} - 3g_{1}^{2}\lambda - 9g_{2}^{2}\lambda + 4\lambda(|y|^{2} + |y'|^{2}) - 2(|y|^{2} + |y'|^{2})^{2} + 12\lambda y_{t}^{2} - 6y_{t}^{4},$$
(3.14)

where λ is defined via the Higgs potential,

$$V = m^2 H^{\dagger} H + \lambda (H^{\dagger} H)^2. \tag{3.15}$$

As can be seen from Eq. (3.14), the stability gives severe constraint for large y and y' regime.

To estimate the stability bound of the Higgs potential, we define the cutoff scale Λ by

$$\lambda(\Lambda) = 0. \tag{3.16}$$

³ The stability without CP violation in the singlet-doublet model is discussed in Refs. [34, 63, 64].

The Lagrangian of the singlet-doublet model is no longer valid above the scale of Λ . Otherwise the lifetime of our vacuum would be shorter than the age of the Universe. We derive the beta function of λ up to the two-loop level by using SARAH [65]. We use the SM beta functions below M_2 , and we change the beta function at $\mu = M_2$. We numerically solve the renormalization group equations at the two-loop level. We calculate the threshold correction at the one-loop level. We find

$$\lambda(M_2+0) - \lambda(M_2-0) \simeq -\frac{1}{3v^2} \left(\Sigma_{1\text{PI}} + 2\lambda v\mathcal{T}\right) + \frac{\lambda}{v^2} \frac{4}{g^2} \Sigma_{WW}^{\text{BSM}}(0), \qquad (3.17)$$

where $\Sigma_{1\text{PI}}$, \mathcal{T} , and Σ_{WW}^{BSM} are given in Appendix A. The details of the derivation are also given in Appendix A. The SM input parameters are given in [66],

$$g'(m_t) = 0.35830, \quad g(m_t) = 0.64779, \quad g_s(m_t) = 1.1666,$$
 (3.18)

$$y_t(m_t) = 0.93690, \quad \lambda(m_t) = 0.12604, \quad m_t = 173.34 \text{ GeV}.$$
 (3.19)

From the analysis of the effective theory, we know that the dimension-5 operator suppressed by $\Lambda = 5$ TeV affects the dark matter annihilation cross section. We require $\Lambda > 50$ TeV (500 TeV) to make the uncertainty coming from the higher dimension operators smaller than 10% (1%).

3.5. Indirect detection

Some region of the parameter space where $0 < \phi < \pi$ can be constrained by DM indirect detection because the DM annihilation process is s wave in the *CP*-violating region. As discussed in Sec. 2, the *CP*-violating interaction term is important to avoid the constraints from the DM direct detection experiments. Thus studying the constraints from indirect detection searches are complementary to the constraints from the direct detection experiments. In the following, we focus on the γ -ray emission by the DM annihilation as in Sec. 2.

3.6. Higgs invisible decay

The Higgs boson can decay into DM pairs if it is heavier than twice the DM mass. This decay mode can be observed as an invisible decay at the LHC. The ATLAS and CMS experiments give the current upper bound on the Higgs invisible branching ratio as $Br(H \rightarrow invisible) < 0.26$ at 95% C.L. [67, 68]. We use this bound for $m_{\rm DM} < m_h/2 \simeq 63$ GeV.



FIG. 5: Current situation (upper panels) and prospects (lower panels) of the singlet-doublet model for $M_2 = 1$ TeV and r = 1. The gray, green, and blue shaded regions are excluded by the XENON1T [2], ACME [40], and Fermi-LAT experiment [37], respectively. The cyan region is excluded by the measurement of the Higgs invisible decay at the LHC experiments [67, 68]. For the prospect, we used LZ [45] and $|d_e| = 10^{-30} e \text{ cm}$ [61, 62].

3.7. Constraints from direct, indirect, eEDM experiments, and stability bound

We show the constraints from the direct and indirect detection experiments, the eEDM experiment, and stability bound. Some prospects are also discussed.

The upper panels in Fig. 5 show the current status of the singlet-doublet model. The constraint

from the XENON1T experiment is very strong. If the model respects CP invariance in the dark sector, only a tiny region around $m_{\rm DM} \sim m_h/2$, the so-called Higgs funnel region, is consistent with the XENON1T experiment. On the other hand, if CP is not a good symmetry in the dark sector, namely $0 < \phi < \pi$, then the wide range of the dark matter mass is possible thanks to the the pseudoscalar interaction associated with the CP violation in the dark sector. Therefore, varying the CP phase is important in the analysis of the singlet-doublet model. The constraint from the ACME experiment excludes 300 GeV $\leq m_{\rm DM} \leq 950$ GeV. For $m_{\rm DM} \geq 950$ GeV, smaller Yukawa couplings can realize the thermal relic abundance thanks to coannihilation processes, and thus eEDM is smaller than the other region. Since the eEDM is proportional to $m_{\rm DM}$, as shown in Eq. (3.13), the lighter mass is less constrained. The constraint from the gamma-ray flux from the dSphs excludes $m_h/2 < m_{\rm DM} < 74$ GeV at 90% confidence level. The constraint from the Higgs invisible decay search excludes $m_{\rm DM} < 54$ GeV at 95% confidence level.

The lower panels in Fig. 5 show the prospect of the model. Here we use the LZ experiment [45] and $|d_e| < 10^{-30} \ e \ cm$ [61, 62] for the prospect of the direct detection experiment and the eEDM measurements, respectively. We find that the *CP*-conserving region can be excluded if there is no signal from the direct detection experiment in future. In the *CP*-conserving case, the DM pair annihilates into $b\bar{b}$ through the Higgs exchange in the s-channel in the Higgs funnel region, and it is p wave. On the other hand, it is s wave in the *CP*-violating case. Therefore, larger Yukawa couplings are required for the *CP*-conserving regions receive the stronger constraint than the *CP*-violating region from the direct detection experiment. The eEDM can cover most of the *CP*-violating region except the Higgs funnel region. In conclusion, if there is no signal from the direct detection experiments nor eEDM, the singlet-doublet model is consistent only at the Higgs funnel region. We address the Higgs funnel region later.

The upper panels in Fig. 6 show the stability bound. We find $\Lambda > 500$ TeV in most of the region of the parameter space. In general, UV physics above the scale Λ gives higher dimensional operators to the Lagrangian of the singlet-doublet model. There is a theoretical uncertainty in this sense, and if $\Lambda = 5$ TeV, then a theoretical uncertainty from the higher dimensional operators in our calculations is not negligible because c_P in Sec. 2 is comparable to $\Lambda^{-1} = (5 \text{ TeV})^{-1}$. If $\Lambda > 500$ TeV, the uncertainty becomes less than 1%. Since the energy density of the dark matter energy density is determined with 1% accuracy [43], we can neglect the uncertainty of ignoring the higher dimension operator from UV physics for $\Lambda > 500$ TeV. Thus, our result shown in Fig. 5 does not receive uncertainty from the higher dimension operators.



FIG. 6: The cutoff scale from the stability bound. The upper (lower) panels are for $M_2 = 1(3)$ TeV and r = 1. In the red, orange, and yellow regions, $\Lambda < 50$ TeV, 50 TeV $\Lambda < 500$ TeV, and 500 TeV Λ , respectively. The gray regions are excluded by the XENON1T experiment [2].

Since the eEDM is proportional to $|M_2|^{-3}$ for $|M_2| \gg |M_1|$, v, we expect that the constraints from the ACME experiment get milder for larger M_2 . We choose $M_2 = 3$ TeV in Fig. 7 and in the lower panels in Fig. 6. We find that the constraint from the eEDM measurement is weak compared with the case for $M_2 = 1$ TeV, as expected. The value of the eEDM for $80 \leq m_{\rm DM} \leq 100$ GeV is below the prospect [61, 62]. However, the stability bound is worse compared to the case for $M_2 = 1$ TeV as can be seen from Fig. 6. In the large region, the cutoff scale is lower than 500 TeV for $M_2 = 3$ TeV, namely the uncertainty by ignoring the higher dimensional operator is



FIG. 7: Current situation (upper panels) and prospects (lower panels) of the singlet-doublet model for $M_2 = 3$ TeV and r = 1. The color notation is the same as Fig. 5.

not negligible. In particular, 50 TeV < Λ < 500 TeV for 80 $\leq m_{\rm DM} \leq$ 100 GeV, and thus the uncertainty is at most 10%. The theoretical prediction in this region highly depends on how we modify the model to push up the cutoff scale.

We conclude that the heavy dark matter region is already excluded by the measurement of the eEDM. Most of the region except for the Higgs funnel region is also excluded if experiments observe null results for the eEDM in future, or we have to accept the $\mathcal{O}(10)\%$ theoretical uncertainty in the calculation of the thermal relic abundance.



FIG. 8: The current and prospects of the direct detection experiments on the Higgs funnel region.

We investigate how to test the model on the Higgs funnel region. First, we focus on the progress of DM direct detection experiments. The Darwin project is an ultimate experiment for the DM direct detection experiment [69]. In Fig. 8, we show the prospect of the LZ and the Darwin in the Higgs funnel region. The spin-independent cross section is highly suppressed in the Higgs funnel region. Nevertheless, we find that the CP-conserving region is excluded by the LZ and Darwin. On the other hand, the large CP-violating region cannot be tested by the direct detection experiment, and thus we need other observables to study the Higgs funnel region completely.

Second, we investigate the Higgs invisible decay. The left panel of Fig. 9 is a magnification of Fig. 5 with the branching ratio of the Higgs decaying into a DM pair. The right panel is the branching ratio for $\phi = 0.5\pi$. If the dark matter mass is very close to a half of the Higgs mass, the branching ratio becomes small due to the phase space suppression. The prospects at future collider experiments on the upper bound on the invisible branching fraction are, for example, 4.8 % at the ILC ($\sqrt{s} = 250 \text{ GeV}$, $\mathcal{L} = 250 \text{ fb}^{-1}$) [70], 2.8 % at the HL-LHC ($\mathcal{L} = 3 \text{ ab}^{-1}$) [71], and 0.5 % at the FCC-ee ($\sqrt{s} = 240 \text{ GeV}$, $\mathcal{L} = 10 \text{ ab}^{-1}$) [72]. Therefore, the region $m_{\text{DM}} \leq 58 \text{ GeV}$ can be probed at the future collider experiments.

Finally, let us discuss the eEDM. Figure 10 shows eEDM in the Higgs funnel region. We find that the eEDM changes rapidly for $0 < \pi < 0.1\pi$ and $0.9\pi < \phi < \pi$. If the dark matter direct detection experiments observe null results in future, the *CP* invariance must be violated in the dark sector as shown in Fig. 8. Consequently, the value of the eEDM must be nonzero if the DM-nucleon



FIG. 9: The branching ratio of the Higgs boson decaying to the DM pairs for $M_2 = 1$ TeV and r = 1. The gray and cyan shaded regions are excluded by the XENON1T experiment [2], and the LHC experiment [67, 68], respectively.

scattering cross section is small enough to avoid the search by the direct detection experiments. Assuming the null results by the direct detection experiments, we find that the minimum value of the eEDM is $d_e = 3 \times 10^{-32} \ e \ cm$ for $|M_2| = 1$ TeV and r = 1 as can be seen from the right panel in Fig. 10. Therefore, we have a chance to test this model by the eEDM measurements in case that experiments reach to $d_e = \mathcal{O}(10^{-32}) \ e \ cm$ in future even if σ_{SI} is very small at the Higgs funnel region. This is a different feature of this model from other Higgs portal DM models, which avoid any constraints from the experiments and are not testable in the Higgs funnel region.

4. CONCLUSION

In this paper, we have investigated the singlet-doublet dark matter model with the CP phase as a UV complete model that predicts the pseudoscalar interaction. Although the scalar sector of the model is the same as the SM and does not contain pseudoscalar fields, the model predicts the pseudoscalar interaction of the DM with the SM Higgs boson because of the CP violation in the dark sector. As a result, the model evades the constraint from the DM direct detection experiments. The DM annihilation processes into the SM particles with the pseudoscalar interaction are s wave, and its cross section at low temperature is not suppressed by powers of velocity. As a result, we can expect the γ -ray emission from dSphs due to the annihilation of DM. Besides, the CP



FIG. 10: The eEDM for $M_2 = 1$ TeV, r = 1. The gray and cyan shaded regions are excluded by the XENON1T experiment [2], and the LHC experiment [67, 68], respectively. The green region in the right panel is excluded by the ACME experiment [40] The blue dashed line is the prospects by [61, 62]. We take $\phi = 0.5\pi$ in the right panel.

violation induces EDMs, which are severely constrained from experiments. Consequently, some of the large CP phase regions are already excluded, though the large CP phase is helpful to avoid the constraint from the direct detection experiments. Therefore, it is essential to investigate the correlation between the direct detection experiments and other observables that are sensitive to CP violation.

We have calculated the γ -ray flux and compared with the result of the Fermi-LAT experiment that gives the constraint of the γ -ray flux from dSphs. We have used 19 dSphs whose J-factors were measured kinematically and listed in the sixth column in Table 1 in Ref. [37]. In this case, we find that the model is excluded for $m_h/2 < m_{\rm DM} < 74$ GeV.

We also have investigated the stability of the Higgs potential. The new fermion fields in the dark sector give negative contributions to the beta function of the Higgs quartic coupling, and thus the stability of the potential becomes worse than the SM case. We have used the beta functions at the two-loop level with the threshold correction at the one-loop level, and calculated the scale where the Higgs quartic coupling becomes 0, $\lambda(\Lambda) = 0$. The analysis of the effective theory for the fermionic DM models implies that the dimension-5 operators with the cutoff scale $\Lambda \leq 500$ TeV affect the calculation of the DM annihilation cross section more than $\mathcal{O}(1)\%$ level. Note that the DM energy density is measured with 1% accuracy. Therefore, our calculation is reliable for $\Lambda > 500$ TeV. With this theoretical constraint, we find that our results do not receive the uncertainty from higher dimensional operators for $m_{\rm DM} < 1$ TeV and $|M_2| < 3$ TeV. Although the model predicts a viable DM candidate without theoretical uncertainty for $\Lambda > 500$ TeV, the Higgs potential is expected to be unstable in the large region of the parameter space because $\Lambda \ll \Lambda_{\rm SM}$ in most of the parameter space, where $\Lambda_{\rm SM} \simeq 10^{10}$ GeV. Therefore, an additional extension of the model is required to make the model valid at a much higher energy scale such as the Planck scale.

We focused on the electron EDM. We found the region where $m_{\rm DM} > 300(2000)$ GeV is excluded for $|M_2| = 1(3)$ TeV. Since the eEDM is proportional to $|M_2|^{-3}$ for $|M_2| \gg |M_1|$, v, we can avoid the constraint from the eEDM with large $|M_2|$. However, the cutoff scale estimated from the stability bound of the Higgs potential becomes lower for the large $|M_2|$ region. Therefore, we conclude that the heavy DM region is already excluded by the measurement of the eEDM. We also found that most of the region except for the Higgs funnel region is also excluded if experiments observe null results for $|d_e| \ge 10^{-30} e$ cm in future, or we have to accept the $\mathcal{O}(10)\%$ theoretical uncertainty in the calculation of the thermal relic abundance.

Finally, we studied the Higgs funnel region where the DM-nucleon scattering cross section and the annihilation cross section of DM are highly suppressed. We find that the *CP*-conserving region can be covered by the LZ and Darwin experiments. For the *CP*-violating region, we cannot expect the direct and indirect detection signals, and thus other observables are required to test the model. The searches for the Higgs invisible decay can probe for the dark matter for $m_{\rm DM} \leq 58$ GeV. Moreover, the eEDM is larger than $3 \times 10^{-32} \ e \ cm$ for $|M_2| = 1$ TeV, r = 1, and $0.1\pi < \phi < 0.9\pi$. If experiments for the eEDM search reach to $\mathcal{O}(10^{-32}) \ e \ cm$, the Higgs invisible decay, and the eEDM.

Acknowledgments

This work was supported by JSPS KAKENHI Grant Number 16K17715 [T.A.].

Appendix A: The threshold correction to the Higgs quartic coupling λ

We determine λ from the on-shell Higgs mass given by

$$m_h^2 = m^2(\mu) + 3\lambda(\mu)v^2(\mu) + (\text{loop corrections}).$$
(A1)

If the theory changes at $\mu = M$, then

$$m^{2}(M-0) + 3\lambda(M-0)v^{2}(M-0) = m^{2}(M+0) + 3\lambda(M+0)v^{2}(M+0) + \Sigma^{\text{BSM}},$$
(A2)

where Σ^{BSM} is loop corrections to the Higgs mass from the new physics sector. From this equation, we can find the threshold correction as follows,

$$\lambda(M+0) - \lambda(M-0) = \lambda(M-0) \left(\frac{v^2(M-0)}{v^2(M+0)} - 1\right) + \frac{m^2(M-0) - m^2(M+0) - \Sigma^{\text{BSM}}}{3v^2(M+0)}$$
$$\simeq \lambda \left(\frac{-\Delta v^2}{v^2}\right) + \frac{-\Delta m^2 - \Sigma^{\text{BSM}}}{3v^2(M+0)},$$
(A3)

where

$$\Delta v^2 = v^2 (M+0) - v^2 (M-0), \tag{A4}$$

$$\Delta m^2 = m^2 (M+0) - m^2 (M-0).$$
(A5)

 Δv^2 is calculated from the Fermi constant that is determined from $\mu \to e \bar{\nu}_e \nu_\mu$. We find

$$\Delta v^2 = -\frac{4}{g^2} \Sigma_{WW}^{BSM}(0) - 2v\mathcal{T}, \tag{A6}$$

where

$$\Sigma_{WW}^{BSM}(0) = -\frac{2}{(4\pi)^2} \Biggl\{ \left(|c_{\chi_j}^L|^2 + |c_{\chi_j}^R|^2 \right) \left(-2B_{00}(0, m_{\chi_j}^2, m_{\chi^{\pm}}^2) + \frac{m_{\chi_j}^2 + m_{\chi^{\pm}}^2}{2} \right) + m_{\chi_j} m_{\chi^{\pm}} \left((c_{\chi_j}^L)^* (c_{\chi_j}^R|^2) + (c_{\chi_j}^L) (c_{\chi_j}^R|^2)^* \right) B_0(0, m_{\chi_j}^2, m_{\chi^{\pm}}^2) \Biggr\}, \quad (A7)$$

$$\mathcal{T} = -\frac{1}{(4\pi)^2} \frac{1}{m_h^2} \sum_{j=1}^3 2g_{\chi_j \chi_j h} m_{\chi_j} A_0(m_{\chi_j}^2).$$
(A8)

We regularize the loop calculation by the $\overline{\text{MS}}$ -scheme, and thus we have to keep calculate tadpole diagrams given in Eq. (A8). The loop functions $(A_0, B_0, \text{ and } B_{00})$ are already regularized by the $\overline{\text{MS}}$ -scheme, and their definitions are given in LoopTools [73]. The couplings are given in Ref. [36]

 Σ^{BSM} is calculated from the loop correction to the Higgs mass from the new physics sector.

$$\Sigma^{\text{BSM}} = \Sigma_{1\text{PI}} + 6\lambda v \mathcal{T},\tag{A9}$$

where

$$\begin{split} \Sigma_{1\mathrm{PI}} &= \frac{2}{(4\pi)^2} \Biggl\{ \sum_{j=1}^3 g_{\chi_j \chi_j h}^2 \left(A_0(m_{\chi_j}^2) + \frac{4m_{\chi_j}^2 - m_h^2}{2} B_0(m_h^2, m_{\chi_j}^2, m_{\chi_j}^2) \right) \\ &+ \sum_{j=1}^3 (g_{\chi_j \chi_j h}^A)^2 \left(A_0(m_{\chi_j}^2) + \frac{-m_h^2}{2} B_0(m_h^2, m_{\chi_j}^2, m_{\chi_j}^2) \right) \\ &+ 2 \sum_{j=1}^3 \sum_{k>j}^3 g_{\chi_j \chi_k h}^2 \left(\frac{A_0(m_{\chi_j}^2) + A_0(m_{\chi_k}^2)}{2} + \frac{(m_{\chi_j} + m_{\chi_k})^2 - m_h^2}{2} B_0(m_h^2, m_{\chi_j}^2, m_{\chi_k}^2) \right) \\ &+ 2 \sum_{j=1}^3 \sum_{k>j}^3 (g_{\chi_j \chi_k h}^A)^2 \left(\frac{A_0(m_{\chi_j}^2) + A_0(m_{\chi_k}^2)}{2} + \frac{(m_{\chi_j} - m_{\chi_k})^2 - m_h^2}{2} B_0(m_h^2, m_{\chi_j}^2, m_{\chi_k}^2) \right) \Biggr\}. \end{split}$$
(A10)

We choose m^2 so as to satisfy that the effective potential has the electroweak vacuum at the one-loop level. The effective potential at the one-loop level is given by

$$\begin{aligned} V_{eff.} &= \frac{m^2}{2} \varphi^2 + \frac{\lambda}{4} \varphi^4 \\ &- \frac{1}{64\pi^2} \left(3(m^2 + \lambda \varphi^2)^2 \left(\frac{3}{2} + \ln \frac{\mu^2}{m^2 + \lambda \varphi^2} \right) \right) \\ &- \frac{1}{64\pi^2} \left((m^2 + 3\lambda \varphi^2)^2 \left(\frac{3}{2} + \ln \frac{\mu^2}{m^2 + 3\lambda \varphi^2} \right) \right) \\ &- \frac{3}{64\pi^2} \left(2 \left(\frac{g^2 \varphi^2}{4} \right)^2 \left(\frac{5}{6} + \ln \frac{\mu^2}{\frac{g^2 \varphi^2}{4}} \right) \right) \\ &- \frac{3}{64\pi^2} \left(\left(\frac{(g^2 + g'^2) \varphi^2}{4} \right)^2 \left(\frac{5}{6} + \ln \frac{\mu^2}{\frac{(g^2 + g'^2) \varphi^2}{4}} \right) \right) \\ &+ \frac{12}{64\pi^2} \left(\left(\frac{y_t^2 \varphi^2}{2} \right)^2 \left(\frac{3}{2} + \ln \frac{\mu^2}{\frac{y_t^2 \varphi^2}{2}} \right) \right) \\ &+ \frac{2}{64\pi^2} \sum_{j=1}^3 \left(m_{\chi_j}^4 \left(\frac{3}{2} + \ln \frac{\mu^2}{m_{\chi_j}^2} \right) \right). \end{aligned}$$
(A11)

We choose m^2 so as to satisfy $V'_{eff.}(v) = 0$ at the one-loop level.

$$m^{2} = -\lambda v^{2} + \frac{1}{(4\pi)^{2}} \Biggl\{ 6\lambda^{2} v^{2} \left(1 + \ln \frac{\mu^{2}}{2\lambda v^{2}} \right) + 6v^{2} \left(\frac{g^{2}}{4} \right)^{2} \left(\frac{1}{3} + \ln \frac{\mu^{2}}{\frac{g^{2} v^{2}}{4}} \right) + 3v^{2} \left(\frac{g^{2} + g^{\prime 2}}{4} \right)^{2} \left(\frac{1}{3} + \ln \frac{\mu^{2}}{\frac{(g^{2} + g^{\prime 2})v^{2}}{4}} \right) - 12v^{2} \left(\frac{y_{t}^{2}}{2} \right)^{2} \left(1 + \ln \frac{\mu^{2}}{\frac{y_{t}^{2} v^{2}}{2}} \right) - 2\frac{1}{v} \sum_{j} g_{\chi_{j}\chi_{j}h} m_{\chi_{j}}^{3} \left(1 + \ln \frac{\mu^{2}}{m_{\chi_{j}}^{2}} \right) \Biggr\}.$$
(A12)

From Eq. (A12), we find

$$\Delta m^2 = \frac{1}{(4\pi)^2} \left[-2\frac{1}{v} \sum_j g_{\chi_j \chi_j h} m_{\chi_j}^3 \left(1 + \ln \frac{\mu^2}{m_{\chi_j}^2} \right) \right].$$
(A13)

Finally, we find

$$\lambda(M+0) - \lambda(M-0) \simeq -\frac{1}{3v^2} \left(\Sigma_{1\text{PI}} + 2\lambda v\mathcal{T}\right) + \frac{\lambda}{v^2} \frac{4}{g^2} \Sigma_{WW}^{\text{BSM}}(0).$$
(A14)

- G. Arcadi, M. Dutra, P. Ghosh, M. Lindner, Y. Mambrini, M. Pierre, S. Profumo and F. S. Queiroz, Eur. Phys. J. C 78, no. 3, 203 (2018) doi:10.1140/epjc/s10052-018-5662-y [arXiv:1703.07364 [hep-ph]].
- [2] E. Aprile *et al.* [XENON Collaboration], Phys. Rev. Lett. **121**, no. 11, 111302 (2018) doi:10.1103/PhysRevLett.121.111302 [arXiv:1805.12562 [astro-ph.CO]].
- [3] X. Cui *et al.* [PandaX-II Collaboration], Phys. Rev. Lett. **119**, no. 18, 181302 (2017) doi:10.1103/PhysRevLett.119.181302 [arXiv:1708.06917 [astro-ph.CO]].
- [4] S. Kanemura, S. Matsumoto, T. Nabeshima and N. Okada, Phys. Rev. D 82, 055026 (2010) doi:10.1103/PhysRevD.82.055026 [arXiv:1005.5651 [hep-ph]].
- [5] L. Lopez-Honorez, T. Schwetz and J. Zupan, Phys. Lett. B 716, 179 (2012) doi:10.1016/j.physletb.2012.07.017 [arXiv:1203.2064 [hep-ph]].
- [6] A. Djouadi, A. Falkowski, Y. Mambrini and J. Quevillon, Eur. Phys. J. C 73, no. 6, 2455 (2013) doi:10.1140/epjc/s10052-013-2455-1 [arXiv:1205.3169 [hep-ph]].
- [7] A. Greljo, J. Julio, J. F. Kamenik, C. Smith and J. Zupan, JHEP 1311, 190 (2013) doi:10.1007/JHEP11(2013)190 [arXiv:1309.3561 [hep-ph]].
- [8] S. Matsumoto, S. Mukhopadhyay and Y. L. S. Tsai, JHEP 1410, 155 (2014) doi:10.1007/JHEP10(2014)155 [arXiv:1407.1859 [hep-ph]].
- [9] A. Beniwal, F. Rajec, C. Savage, P. Scott, C. Weniger, M. White and A. G. Williams, Phys. Rev. D 93, no. 11, 115016 (2016) doi:10.1103/PhysRevD.93.115016 [arXiv:1512.06458 [hep-ph]].
- [10] S. Matsumoto, S. Mukhopadhyay and Y. L. S. Tsai, Phys. Rev. D 94, no. 6, 065034 (2016) doi:10.1103/PhysRevD.94.065034 [arXiv:1604.02230 [hep-ph]].
- [11] P. Athron *et al.* [GAMBIT Collaboration], arXiv:1808.10465 [hep-ph].
- [12] S. Baum, M. Carena, N. R. Shah and C. E. M. Wagner, JHEP **1804**, 069 (2018) doi:10.1007/JHEP04(2018)069 [arXiv:1712.09873 [hep-ph]].
- [13] R. Mahbubani and L. Senatore, Phys. Rev. D 73, 043510 (2006) doi:10.1103/PhysRevD.73.043510
 [hep-ph/0510064].
- [14] F. D'Eramo, Phys. Rev. D 76, 083522 (2007) doi:10.1103/PhysRevD.76.083522 [arXiv:0705.4493 [hep-ph]].

- [15] R. Enberg, P. J. Fox, L. J. Hall, A. Y. Papaioannou and M. Papucci, JHEP 0711, 014 (2007) doi:10.1088/1126-6708/2007/11/014 [arXiv:0706.0918 [hep-ph]].
- [16] J. L. Feng, K. T. Matchev and F. Wilczek, Phys. Lett. B 482, 388 (2000) doi:10.1016/S0370-2693(00)00512-8 [hep-ph/0004043].
- [17] A. Pierce, Phys. Rev. D 70, 075006 (2004) doi:10.1103/PhysRevD.70.075006 [hep-ph/0406144].
- [18] N. Arkani-Hamed, A. Delgado and G. F. Giudice, Nucl. Phys. B 741, 108 (2006) doi:10.1016/j.nuclphysb.2006.02.010 [hep-ph/0601041].
- [19] K. Hamaguchi and K. Ishikawa, Phys. Rev. D 93, no. 5, 055009 (2016) doi:10.1103/PhysRevD.93.055009
 [arXiv:1510.05378 [hep-ph]].
- [20] T. Cohen, J. Kearney, A. Pierce and D. Tucker-Smith, Phys. Rev. D 85, 075003 (2012) doi:10.1103/PhysRevD.85.075003 [arXiv:1109.2604 [hep-ph]].
- [21] C. Cheung, L. J. Hall, D. Pinner and J. T. Ruderman, JHEP 1305, 100 (2013) doi:10.1007/JHEP05(2013)100 [arXiv:1211.4873 [hep-ph]].
- [22] C. Cheung and D. Sanford, JCAP 1402, 011 (2014) doi:10.1088/1475-7516/2014/02/011
 [arXiv:1311.5896 [hep-ph]].
- [23] L. Calibbi, A. Mariotti and P. Tziveloglou, JHEP 1510, 116 (2015) doi:10.1007/JHEP10(2015)116
 [arXiv:1505.03867 [hep-ph]].
- [24] A. Freitas, S. Westhoff and J. Zupan, JHEP 1509, 015 (2015) doi:10.1007/JHEP09(2015)015
 [arXiv:1506.04149 [hep-ph]].
- [25] G. Cynolter, J. Kovács and E. Lendvai, Mod. Phys. Lett. A **31**, no. 01, 1650013 (2016) doi:10.1142/S0217732316500139 [arXiv:1509.05323 [hep-ph]].
- [26] M. Badziak, M. Olechowski and P. Szczerbiak, JHEP 1603, 179 (2016) doi:10.1007/JHEP03(2016)179
 [arXiv:1512.02472 [hep-ph]].
- [27] S. Horiuchi, O. Macias, D. Restrepo, A. Rivera, O. Zapata and H. Silverwood, JCAP 1603, no. 03, 048 (2016) doi:10.1088/1475-7516/2016/03/048 [arXiv:1602.04788 [hep-ph]].
- [28] S. Banerjee, S. Matsumoto, K. Mukaida and Y. L. S. Tsai, JHEP 1611, 070 (2016) doi:10.1007/JHEP11(2016)070 [arXiv:1603.07387 [hep-ph]].
- [29] C. Cai, Z. H. Yu and H. H. Zhang, Nucl. Phys. B 921, 181 (2017) doi:10.1016/j.nuclphysb.2017.05.015 [arXiv:1611.02186 [hep-ph]].
- [30] J. Kearney, N. Orlofsky and A. Pierce, Phys. Rev. D 95, no. 3, 035020 (2017) doi:10.1103/PhysRevD.95.035020 [arXiv:1611.05048 [hep-ph]].
- [31] P. Huang, R. A. Roglans, D. D. Spiegel, Y. Sun and C. E. M. Wagner, Phys. Rev. D 95, no. 9, 095021 (2017) doi:10.1103/PhysRevD.95.095021 [arXiv:1701.02737 [hep-ph]].
- [32] M. Badziak, M. Olechowski and P. Szczerbiak, Phys. Lett. B 770, 226 (2017) doi:10.1016/j.physletb.2017.04.059 [arXiv:1701.05869 [hep-ph]].
- [33] Q. F. Xiang, X. J. Bi, P. F. Yin and Z. H. Yu, Phys. Rev. D 97, no. 5, 055004 (2018) doi:10.1103/PhysRevD.97.055004 [arXiv:1707.03094 [hep-ph]].

- [34] A. Voigt and S. Westhoff, JHEP 1711, 009 (2017) doi:10.1007/JHEP11(2017)009 [arXiv:1708.01614 [hep-ph]].
- [35] T. Abe, R. Kitano and R. Sato, Phys. Rev. D 91, no. 9, 095004 (2015) Erratum: [Phys. Rev. D 96, no. 1, 019902 (2017)] doi:10.1103/PhysRevD.96.019902, 10.1103/PhysRevD.91.095004 [arXiv:1411.1335 [hep-ph]].
- [36] T. Abe, Phys. Lett. B 771, 125 (2017) doi:10.1016/j.physletb.2017.05.048 [arXiv:1702.07236 [hep-ph]].
- [37] A. Albert *et al.* [Fermi-LAT and DES Collaborations], Astrophys. J. **834**, no. 2, 110 (2017) doi:10.3847/1538-4357/834/2/110 [arXiv:1611.03184 [astro-ph.HE]].
- [38] M. Pospelov and A. Ritz, Annals Phys. 318, 119 (2005) doi:10.1016/j.aop.2005.04.002 [hep-ph/0504231].
- [39] M. E. Pospelov and I. B. Khriplovich, Sov. J. Nucl. Phys. 53, 638 (1991) [Yad. Fiz. 53, 1030 (1991)].
- [40] V. Andreev et al. [ACME Collaboration], Nature 562, no. 7727, 355 (2018). doi:10.1038/s41586-018-0599-8
- [41] S. Gopalakrishna and A. Velusamy, arXiv:1812.11303 [hep-ph].
- [42] G. Bélanger, F. Boudjema, A. Pukhov and A. Semenov, Comput. Phys. Commun. 192, 322 (2015) doi:10.1016/j.cpc.2015.03.003 [arXiv:1407.6129 [hep-ph]].
- [43] P. A. R. Ade *et al.* [Planck Collaboration], Astron. Astrophys. **594**, A13 (2016) doi:10.1051/0004-6361/201525830 [arXiv:1502.01589 [astro-ph.CO]].
- [44] E. Aprile et al. [XENON Collaboration], JCAP 1604, no. 04, 027 (2016) doi:10.1088/1475-7516/2016/04/027 [arXiv:1512.07501 [physics.ins-det]].
- [45] D. S. Akerib et al. [LUX-ZEPLIN Collaboration], arXiv:1802.06039 [astro-ph.IM].
- [46] K. Ghorbani, JCAP 1501, 015 (2015) doi:10.1088/1475-7516/2015/01/015 [arXiv:1408.4929 [hep-ph]].
- [47] S. Baek, P. Ko and J. Li, Phys. Rev. D 95, no. 7, 075011 (2017) doi:10.1103/PhysRevD.95.075011 [arXiv:1701.04131 [hep-ph]].
- [48] K. Ghorbani and P. H. Ghorbani, arXiv:1812.04092 [hep-ph].
- [49] S. Ipek, D. McKeen and A. E. Nelson, Phys. Rev. D 90, no. 5, 055021 (2014) doi:10.1103/PhysRevD.90.055021 [arXiv:1404.3716 [hep-ph]].
- [50] J. M. No, Phys. Rev. D 93, no. 3, 031701 (2016) doi:10.1103/PhysRevD.93.031701 [arXiv:1509.01110 [hep-ph]].
- [51] D. Goncalves, P. A. N. Machado and J. M. No, Phys. Rev. D 95, no. 5, 055027 (2017) doi:10.1103/PhysRevD.95.055027 [arXiv:1611.04593 [hep-ph]].
- [52] M. Bauer, U. Haisch and F. Kahlhoefer, JHEP **1705**, 138 (2017) doi:10.1007/JHEP05(2017)138
 [arXiv:1701.07427 [hep-ph]].
- [53] P. Tunney, J. M. No and M. Fairbairn, Phys. Rev. D 96, no. 9, 095020 (2017) doi:10.1103/PhysRevD.96.095020 [arXiv:1705.09670 [hep-ph]].
- [54] G. Arcadi, M. Lindner, F. S. Queiroz, W. Rodejohann and S. Vogl, JCAP 1803, no. 03, 042 (2018) doi:10.1088/1475-7516/2018/03/042 [arXiv:1711.02110 [hep-ph]].

- [55] P. Pani and G. Polesello, Phys. Dark Univ. 21, 8 (2018) doi:10.1016/j.dark.2018.04.006 [arXiv:1712.03874 [hep-ph]].
- [56] N. F. Bell, G. Busoni and I. W. Sanderson, JCAP 1808, no. 08, 017 (2018) doi:10.1088/1475-7516/2018/08/017 [arXiv:1803.01574 [hep-ph]].
- [57] T. Li, Phys. Lett. B 782, 497 (2018) doi:10.1016/j.physletb.2018.05.073 [arXiv:1804.02120 [hep-ph]].
- [58] T. Abe, M. Fujiwara and J. Hisano, arXiv:1810.01039 [hep-ph].
- [59] T. Abe et al. [LHC Dark Matter Working Group], arXiv:1810.09420 [hep-ex].
- [60] G. Bélanger, F. Boudjema, A. Pukhov and A. Semenov, Comput. Phys. Commun. 192, 322 (2015) doi:10.1016/j.cpc.2015.03.003 [arXiv:1407.6129 [hep-ph]].
- [61] D. M. Kara, I. J. Smallman, J. J. Hudson, B. E. Sauer, M. R. Tarbutt and E. A. Hinds, New J. Phys. 14, 103051 (2012) doi:10.1088/1367-2630/14/10/103051 [arXiv:1208.4507 [physics.atom-ph]].
- [62] D. Kawall, J. Phys. Conf. Ser. 295, 012031 (2011). doi:10.1088/1742-6596/295/1/012031
- [63] C. Cheung, M. Papucci and K. M. Zurek, JHEP **1207**, 105 (2012) doi:10.1007/JHEP07(2012)105 [arXiv:1203.5106 [hep-ph]].
- [64] J. W. Wang, X. J. Bi, P. F. Yin and Z. H. Yu, arXiv:1811.08743 [hep-ph].
- [65] F. Staub, arXiv:0806.0538 [hep-ph].
- [66] D. Buttazzo, G. Degrassi, P. P. Giardino, G. F. Giudice, F. Sala, A. Salvio and A. Strumia, JHEP 1312, 089 (2013) doi:10.1007/JHEP12(2013)089 [arXiv:1307.3536 [hep-ph]].
- [67] The ATLAS collaboration [ATLAS Collaboration], ATLAS-CONF-2018-054.
- [68] A. M. Sirunyan et al. [CMS Collaboration], [arXiv:1809.05937 [hep-ex]].
- [69] J. Aalbers et al. [DARWIN Collaboration], JCAP 1611, 017 (2016) doi:10.1088/1475-7516/2016/11/017
 [arXiv:1606.07001 [astro-ph.IM]].
- [70] D. M. Asner *et al.*, arXiv:1310.0763 [hep-ph].
- [71] CMS Collaboration [CMS Collaboration], CMS-PAS-HIG-16-016.
- [72] M. Bicer et al. [TLEP Design Study Working Group], JHEP 1401, 164 (2014) doi:10.1007/JHEP01(2014)164 [arXiv:1308.6176 [hep-ex]].
- [73] T. Hahn and M. Perez-Victoria, Comput. Phys. Commun. 118, 153 (1999) doi:10.1016/S0010-4655(98)00173-8 [hep-ph/9807565].