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**ACROPOLIS: A Generic Framework for
Photodisintegration Of Light Elements**

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ACROPOLIS: A **generic framework for Photodisintegration Of Light elements**

Paul Frederik Depta^a, Marco Hufnagel^{a b}, and Kai Schmidt-Hoberg^a

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ABSTRACT:

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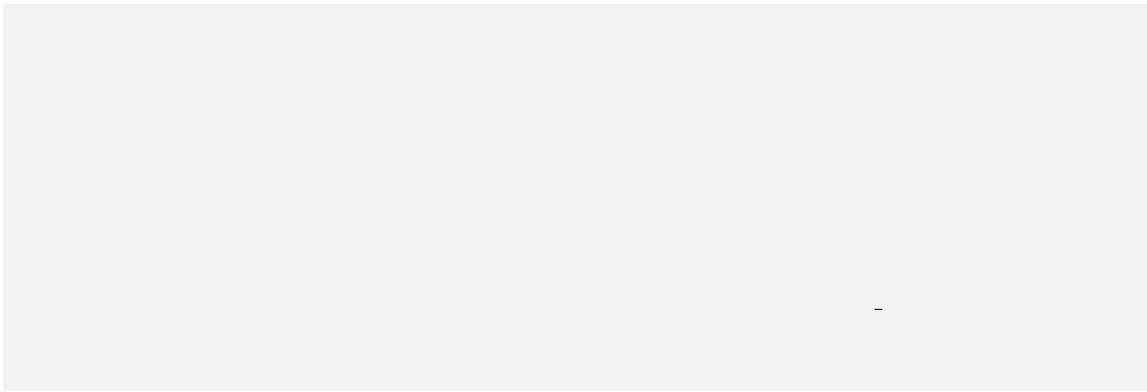
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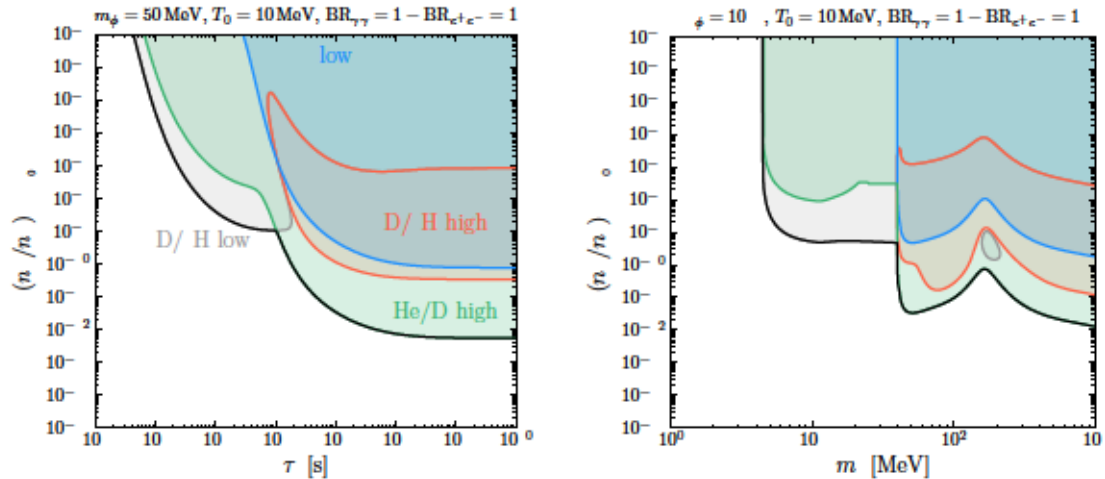


Figure 1. 95% C.L. constraints for decay of a decoupled MeV-scale BSM particle (implemented in `DecayModel`) into two photons ($\text{BR}_{\gamma\gamma} = 1 - \text{BR}_{e^+e^-} = 1$) in $\tau_\phi - (n_\phi/n_\gamma)|_{T=T_0}$ plane (left) and $m_\phi - (n_\phi/n_\gamma)|_{T=T_0}$ plane (right) with $T_0 = 10$ MeV and $m_\phi = 50$ MeV (left) as well as $\tau_\phi = 10^7$ s (right). The limits from individual observables are shown separately: primordial deuterium abundance (orange high, grey low), helium-4 mass fraction \mathcal{Y}_p (blue), and helium-3 abundance normalised by deuterium (green). The overall 95% C.L. BBN limit is given by the black full line as an envelope of individual 95% C.L. constraints neglecting correlations. Using $(n_\phi/n_\gamma)|_{T=T_0}$, i.e. `n0a`, as a **fast** parameter on a single computing node with two AMD EPYC 7402 24-Core Processors the scans took ~ 40 min (left) and ~ 2 h (right) for a 200×200 grid.

deuterium, $m_\phi = 2E_D^{\text{th}} \approx 4.4$ MeV. Apart from some regions with more complex structure due to different disintegration reactions the limits become increasingly strong with larger m_ϕ as the energy density injected into the SM becomes larger.

The scans for figure 1 took ~ 40 min (left) and ~ 2 h (right) for a 200×200 grid on an AMD EPYC 7402 24-Core Processors, clearly highlighting the performance improvement due to the **fast** parameter $(n_\phi/n_\gamma)|_{T=T_0}$, i.e. `n0a`, making the number of points in this direction computationally inexpensive (cf. also appendix B). The runtime is thus determined mostly by the number of points in the direction of τ_ϕ or m_ϕ (not **fast**). Note that the longer runtime for the right panel is a result of the database files for the electromagnetic cascade reaction rates having an upper limit on the energy of $m_\phi/2 = E_0 = 100$ MeV, which often corresponds to the most interesting region in parameter space. For masses above the pion threshold in particular, $m_\phi \gtrsim 280$ MeV, hadrodisintegration may become relevant if ϕ has non-vanishing couplings to quarks, implying that $\text{BR}_{\gamma\gamma} + \text{BR}_{e^+e^-} < 1$ in general. Also muons are kinematically available in the mass region (which are currently not implemented in `ACROPOLIS`).

In figure 2 we show the constraints for residual annihilations of DM into two photons ($\text{BR}_{\gamma\gamma} = 1 - \text{BR}_{e^+e^-} = 1$) for purely *s*-wave annihilations (left, $b = 0$) and purely *p*-wave annihilations (right, $a = 0$, $T_{\text{kd}} = 1$ MeV) as implemented in `AnnihilationModel`. These limits start at the disintegration threshold of deuterium, $m_\chi = E_D^{\text{th}} \approx 2.2$ MeV, and closely

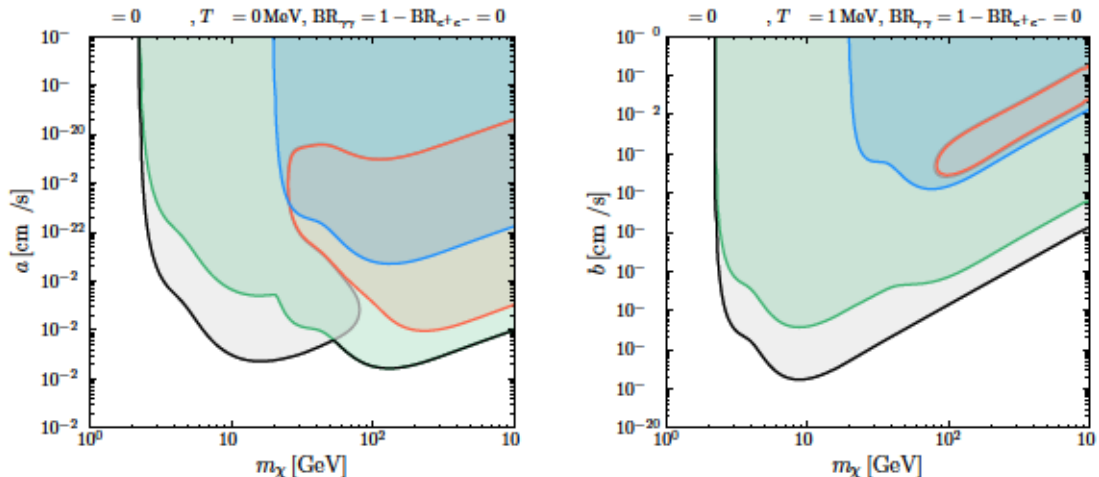


Figure 2. 95% C.L. constraints for residual annihilations of DM (implemented in `AnnihilationModel`) into two photons ($\text{BR}_{\gamma\gamma} = 1 - \text{BR}_{e^+e^-} = 1$) for purely s -wave annihilations (left, $b = 0$) and purely p -wave annihilations (right, $a = 0$, $T_{\text{kd}} = 1$ MeV). For the explanation of the colour-coding see figure 1. Using $(n_\phi/n_\gamma)|_{T=T_b}$, i.e. `n0a`, as a fast parameter on a single computing node with two AMD EPYC 7402 24-Core Processors the scans took ~ 10 h each for a 200×200 grid.

resemble those presented in [26], albeit with updated observationally inferred primordial abundances. We therefore refer to [26] for a detailed discussion. The scans took ~ 10 h each for a 200×200 grid on the aforementioned computing node.

6 Implementing your own models

6.1 The model framework `acropolis.models`

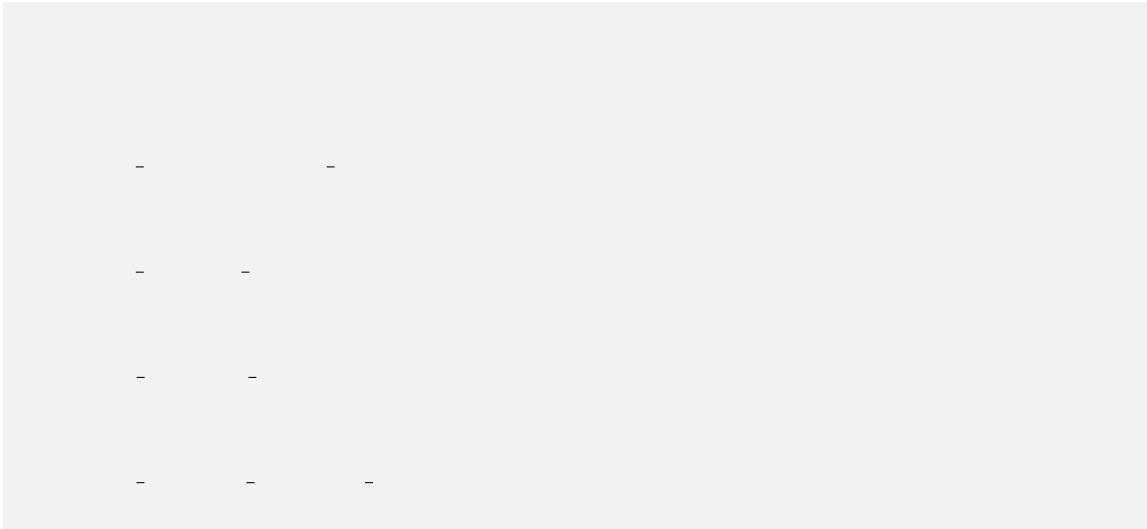
While the provided example models should suffice to tackle most problems of interest, it may sometimes still happen that a scenario cannot directly be mapped to the standard implementation in `ACROPOLIS`. For such cases, `ACROPOLIS` provides further tools that allow for an easy implementation of additional models. The most important class in this context is `acropolis.models.AbstractModel`, which is an abstract base class containing most of the low-level implementation needed to run its method `run_disintegration()`. In fact, using this class as a base, any new model can be implemented in only two steps:

- (i) create a new class, say `NewModel`, that uses `AbstractModel` as a base class, and
- (ii) implement all abstract methods that are provided by `AbstractModel`, i.e.⁷

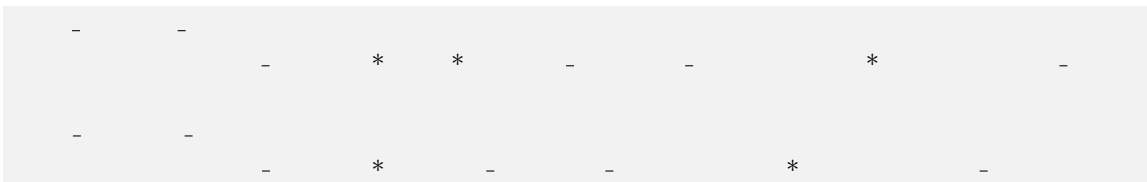
- `AbstractModel._temperature_range()`

⁷By default, the function `AbstractModel._source_positron()` simply returns the output of `AbstractModel._source_electron()`, which is justified for most scenarios. However, if your specific scenario predicts different source terms for electrons and positrons, it is always possible to simply overwrite the former function.

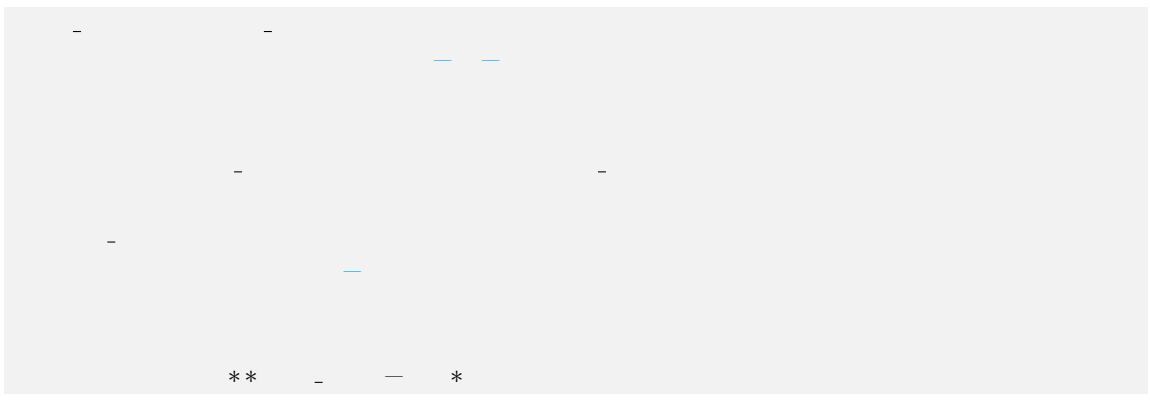
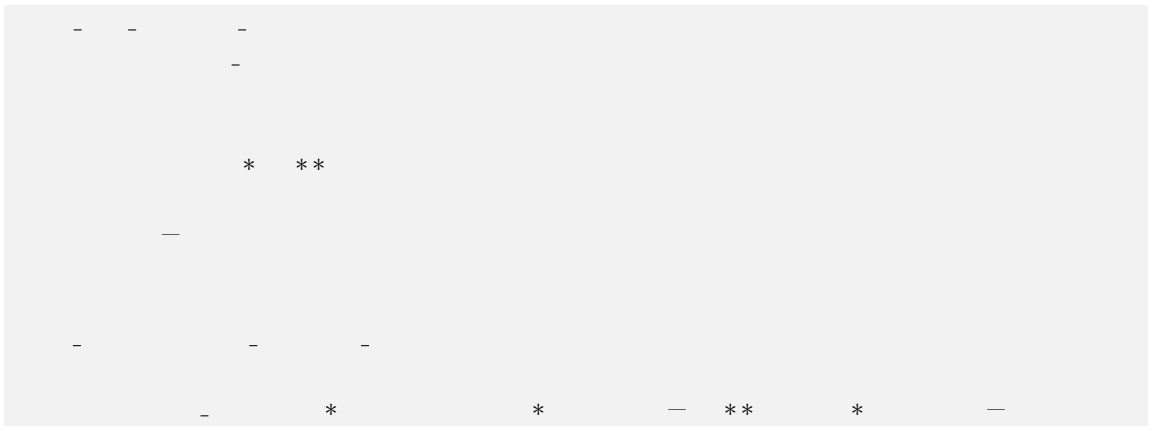
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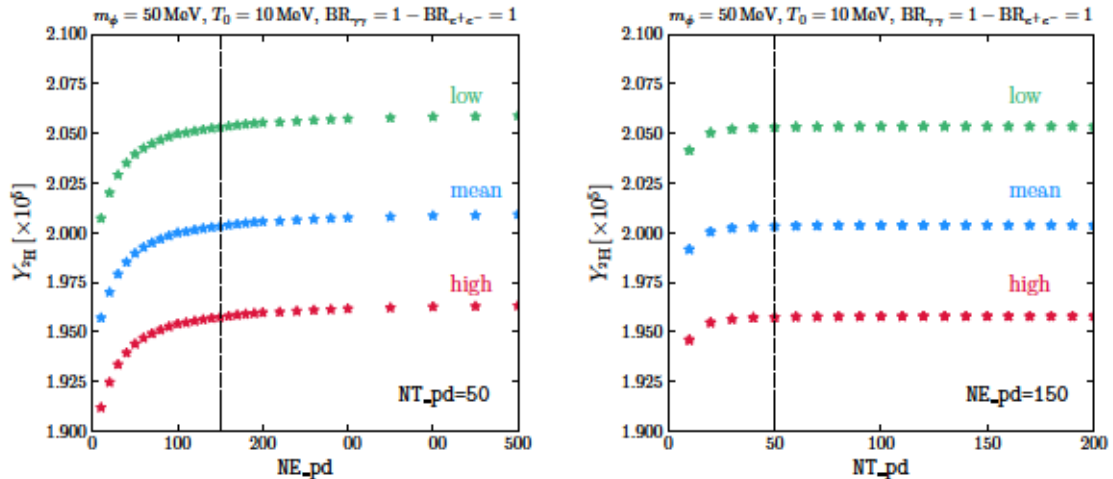


Figure 3. Convergence of the abundance of deuterium as a function of the grid points NE_pd and NT_pd . The dashed line indicates the default value in ACROPOLIS.

Acknowledgments

This work is supported by the ERC Starting Grant ‘NewAve’ (638528), the Deutsche Forschungsgemeinschaft under Germany’s Excellence Strategy – EXC 2121 ‘Quantum Universe’ – 390833306, and by the F.R.S. FNRS under the Excellence of Science (EoS) project No. 30820817 be.h ‘The H boson gateway to physics beyond the Standard Model’.

A Rates for the cascade processes

In this appendix, we collect for completeness all relevant total and differential interaction rates $\Gamma_x(E)$ and $K_{x' \rightarrow x}(E, E')$ for the cascade processes of high-energetic photons, electrons, and positrons on the background photons, electrons, and nuclei (see eqs. (2.4) and (2.7)). Large parts are directly taken from [24].

Target densities

The thermal photon spectrum differential in energy $f_\gamma(\bar{\epsilon})$ is given by

$$f_\gamma(\bar{\epsilon}) = \frac{\bar{\epsilon}^2}{\pi^2} \times \frac{1}{\exp(\bar{\epsilon}/T) - 1}, \quad (\text{A.1})$$

while the total baryon number density can be calculated from the baryon-to-photon ratio η and the number density of photons $n_\gamma(T)$,

$$n_b(T) = \eta \times n_\gamma(T) = \eta \times \frac{2\zeta(3)}{\pi^2} T^3. \quad (\text{A.2})$$

Via charge neutrality we obtain for the number density of background electrons

$$n_e(T) = \sum_N Z_N n_N \simeq [Y_p(T) + 2Y_{4\text{He}}(T)] \times n_b(T), \quad Y_N(T) = \frac{n_N(T)}{n_b(T)}. \quad (\text{A.3})$$

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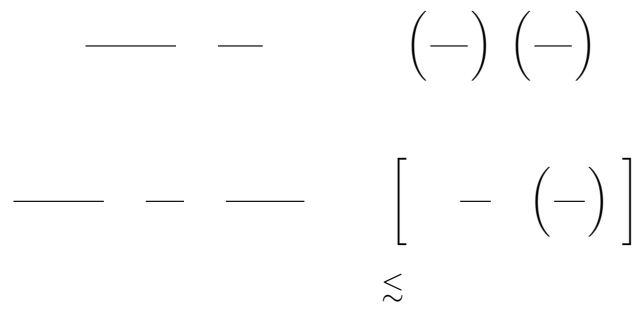
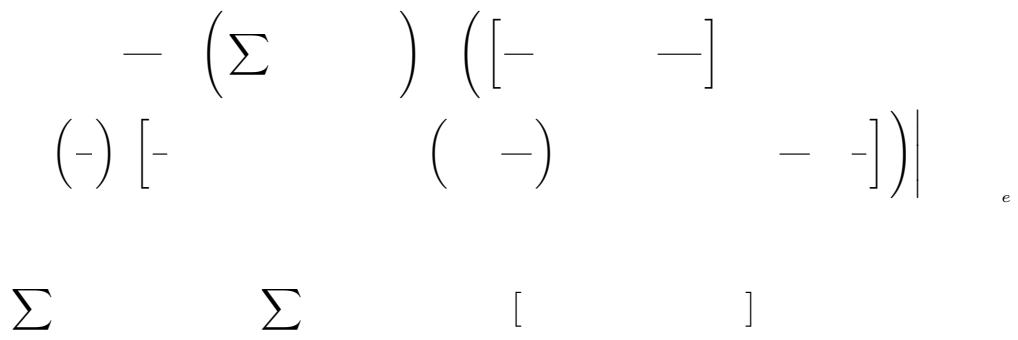
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