

# Commissioning of TREX-DM, a Low Background Micromegas-based Time Projection Chamber for Low Mass WIMP Detection

*F. J. Iguaz\**, *J. García Garza*, *F. Aznar†*, *J. F. Castel*, *S. Cebrián*, *T. Dafni*, *J. A. García*,  
*I. G. Irastorza*, *A. Lagraba*, *G. Luzón*, *A. Peiró*

Laboratorio de Física Nuclear y Astropartículas, Universidad de Zaragoza, Spain

DOI: [http://dx.doi.org/10.3204/DESY-PROC-2015-02/iguaz\\_francisco](http://dx.doi.org/10.3204/DESY-PROC-2015-02/iguaz_francisco)

Dark Matter experiments are recently focusing their detection techniques in low-mass WIMPs, something which requires the use of light elements and low energy threshold. In this context, we describe the TREX-DM experiment, a low background Micromegas-based Time Projection Chamber for low-mass WIMP detection. Its main goal is the operation of an active detection mass  $\sim 0.3$  kg, with an energy threshold below 0.4 keVee and fully built with previously selected radiopure materials. This work focuses on the commissioning of the actual setup situated in a laboratory on surface. A preliminary background model of the experiment is also presented, based on Geant4 simulations and two discrimination methods: a conservative muon/electron and one based on a  $^{252}\text{Cf}$  source. Based on this model, TREX-DM could be competitive in the search for low mass WIMPs and, in particular, it could be sensitive to the WIMP interpretation of the DAMA/LIBRA hint.

## 1 Motivation

The main strategy of Dark Matter experiments [1] is based on accumulating large target masses of heavy nuclei (like Xenon), keeping low background levels by a systematic radiopurity control of all components and an enhancement of the electron/neutron discrimination methods. However, some recent positive hints, which may be interpreted in terms of low mass WIMPs, have changed the detection strategy to sub-keV energies and light gases. This research line could be led in future experiments by Time Projection Chambers (TPCs), as they can reach energy thresholds  $\sim 100$  eV and have access to richer topological information. In contrast to current gaseous-based experiments, focused on directional Dark Matter detection [2], the TREX-DM experiment proposes a strategy based on high gas pressures, even if neutron/electron discrimination could be less effective, but keeping a low energy threshold. TREX-DM is a low background Micromegas-based TPC for low-mass WIMP detection and will profit from all developments made in Micromegas technology [3, 4], as well as in the selection of radiopure materials [5, 6], specially in CAST [7] and NEXT-MM [8] projects. Its main goal is the operation of an active detection mass  $\sim 0.3$  kg with an energy threshold below 0.4 keVee (as already observed in [7]).

---

\*Corresponding author (iguaz@unizar.es)

†Present address: Centro Universitario de la Defensa, Universidad de Zaragoza, Spain.



### 3 Background model of TREX-DM

The sensitivity of the experiment has been studied creating a first background model which reproduces the conditions at the Canfranc Underground Laboratory (LSC). We have considered two light gas mixtures at 10 bar: Ar+2% $i$ C<sub>4</sub>H<sub>10</sub> and Ne+2% $i$ C<sub>4</sub>H<sub>10</sub>, with an active mass of 0.3 and 0.16 kg respectively and which are good candidates to detect low mass WIMPs. However, the sensitivity of an argon-based mixture may be limited by one of its isotopes ( $^{39}$ Ar), which is  $\beta$ -emitter and has a long life-time. In our model, we have considered the lowest content of this isotope, measured in argon extracted from underground sources [10]. We have also simulated the main radioactive isotopes of all the inner components using their measured activities [5, 6] and the cosmic muon flux in Canfranc. In some cases like the Micromegas detectors or their connectors, we have considered the activities of their radiopure alternative. The external gamma flux has not been included as its contribution may be suppressed by an external shielding.

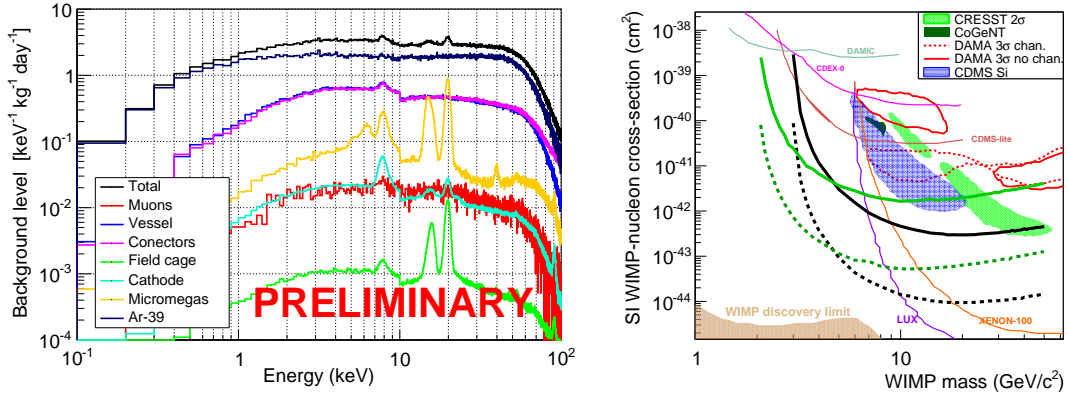


Figure 2: Left: Background spectrum expected in TREX-DM experiment (black line) during a physics run in an underground laboratory if operated in Ar+2% $i$ C<sub>4</sub>H<sub>10</sub> at 10 bar. The contribution of the different simulated components is also plotted: external muon flux (red line), vessel contamination (blue line), connectors (magenta line), field cage (green line), central cathode (brown line), Micromegas detector (purple line) and  $^{39}$ Ar (dark blue line). Right: WIMP parameter space focused on the low-mass range. Filled regions represent the values that may explain the hints of positive signals observed in CoGeNT, CDMS-Si, CRESST and DAMA/LIBRA experiments. The thick lines are the preliminary sensitivity of TREX-DM surpassing a 0.4 keVee energy threshold and two different hypothesis on background and exposure: 100 (solid) and 1(dashed) keV<sup>-1</sup> kg<sup>-1</sup> day<sup>-1</sup>, and 1 and 10 kg-year respectively, and for both argon- (black) and neon-based (green) mixtures.

Two analysis have been used in this background model. The first one is a modified version of the CAST one [7], optimized to discriminate low energy X-rays from complex topologies like gammas and cosmic muons. It uses two likelihood functions generated by the X-rays' cluster features of a calibration source. Fixing a total of 80% signal efficiency, the expected background level for an argon- (neon-) based mixture gas at 10 bar is  $\sim 3.1$  ( $\sim 1.4$ ) keV<sup>-1</sup> kg<sup>-1</sup> day<sup>-1</sup>, dominated by the  $^{39}$ Ar isotope in the case of argon and by the connectors and the vessel in the case of neon. The contribution of each component is shown in Fig. 2 (left) for the argon

case. The second analysis is based on the simulation of a  $^{252}\text{Cf}$  neutron source, which reproduces better WIMPs signals. The level obtained in argon is a  $\sim 44\%$  lower, as nuclear recoils show narrower cluster widths. Supposing a 0.4 keVee energy threshold and former background levels, the TREX-DM experiment could be sensitive to a relevant fraction of the low-mass WIMP parameter space (see Fig. 2, right) including the regions invoked in some interpretations of the DAMA/LIBRA results and other hints of positive WIMPs signals, with an exposure of 1 kg-year.

## 4 Conclusions and prospects

TREX-DM is a low background Micromegas-based TPC for low-mass WIMP detection. Its main goal is the operation of a light gas at high pressure (active mass  $\sim 0.3$  kg) with an energy threshold of 0.4 keVee or below and fully built with previously selected radiopure materials. The detector is being commissioned at TREX laboratory and may be installed at the LSC during 2016 for a possible physics run.

## Acknowledgments

We acknowledge the Micromegas workshop of IRFU/SEDI and the Servicio General de Apoyo a la Investigación-SAI of the University of Zaragoza. We acknowledge the support from the European Commission under the European Research Council T-REX Starting Grant ref. ERC-2009-StG-240054 of the IDEAS program of the 7th EU Framework Program. We also acknowledge support from the Spanish Ministry MINECO under contracts ref. FPA2008-03456 and FPA2011-24058, as well as under the CPAN project ref. CSD2007-00042 from the Consolider-Ingenio 2010 program. These grants are partially funded by the European Regional Development funded (ERDF/FEDER). F.I. acknowledges the support from the Juan de la Cierva program and T.D. from the Ramón y Cajal program of MICINN.

## References

- [1] L. Baudis, “Direct dark matter detection: The next decade”, *Physics of the Dark Universe* **1**, 94 (2012).
- [2] S. Ahlen et al., “The Case for a Directional Dark Matter Detector and the Status of Current Experimental Efforts”, *Int. Jour. Mod. Phys. A* **25**, 1 (2010).
- [3] I. Giomataris et al. “Micromegas in a bulk”, *Nucl. Instrum. Meth. A* **560**, 405 (2006).
- [4] S. Andriamonje, D. Attie, E. Berthoumieux, M. Calviani, P. Colas et al., “Development and performance of Microbulk Micromegas detectors”, *JINST* **5**, P02001 (2010).
- [5] S. Cebrián et al., “Radiopurity of micromegas readout planes”, *Astropart. Phys.* **34**, 354 (2011).
- [6] F. Aznar et al. “Assesment of material radiopurity for Rare Event experiments using Micromegas”, *JINST* **8**, C11012 (2013).
- [7] S. Aune, J. Castel, T. Dafni, M. Davenport, G. Fanourakis et al., “Low background x-ray detection with Micromegas for axion search”, *JINST* **9**, P01001 (2014).
- [8] V. Alvarez et al. “Description and commissioning of NEXT-MM prototype: first results from operation in a Xenon-Trimethylamine gas mixture”, *JINST* **9**, P03010 (2014).
- [9] F.J. Iguaz et al., “Micromegas detector developments for Dark Matter directional detection with MIMAC”, *JINST* **6**, P07002 (2011).
- [10] J. Xu et al., “A study of the trace  $^{39}\text{Ar}$  content in argon from deep underground sources”, *Astrop. Part.* **66**, 53 (2015).