



## Erratum: “Jupiter’s Composition Suggests Its Core Assembled Exterior to the N<sub>2</sub> Snowline” (2019, AJ, 158, 194)

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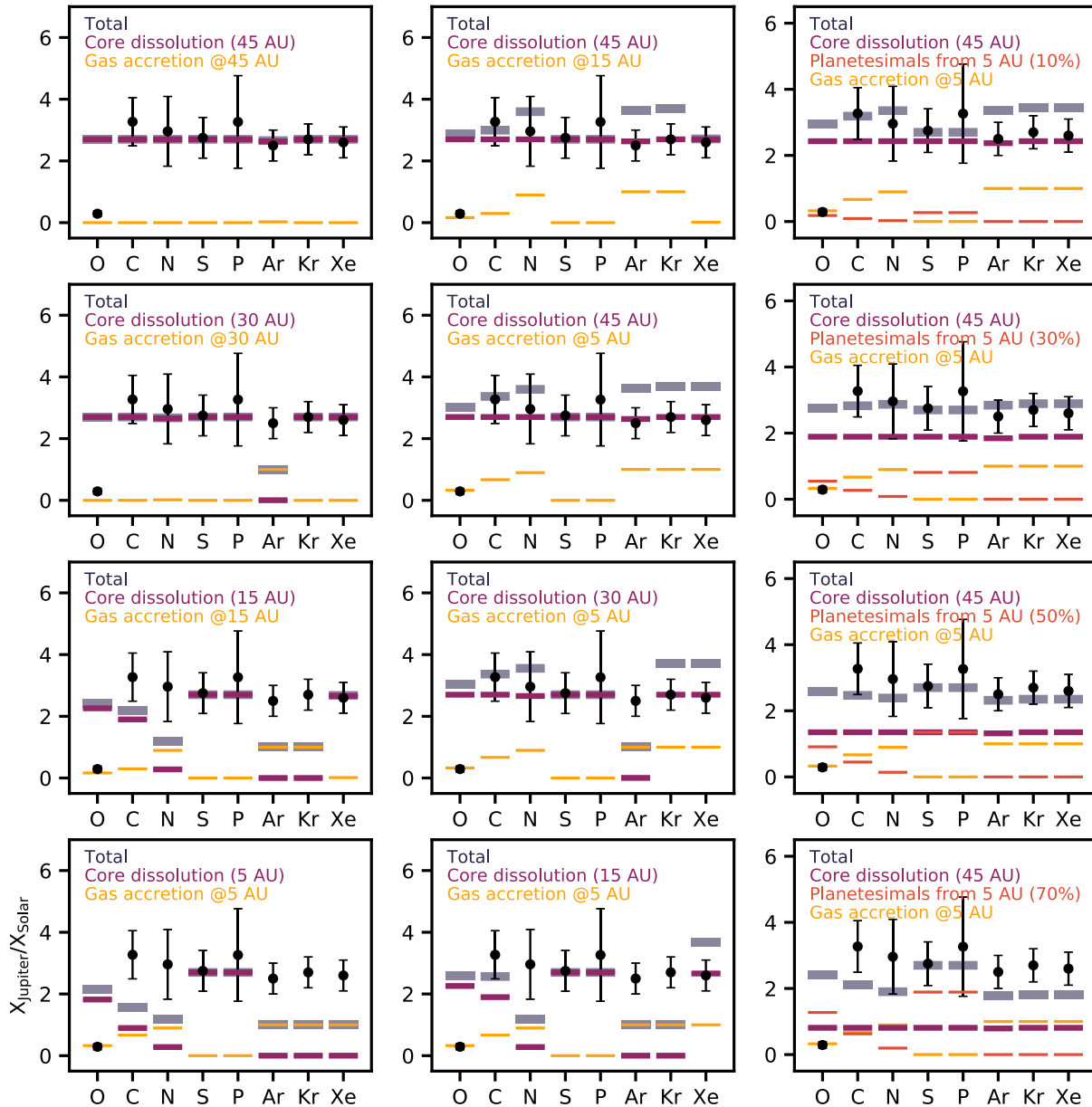
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In the published article, there was a mismatch between the legend and line colors for the gas and planetesimal accretion contributions to Jupiter’s O, C, N, S, P, Ar, Kr, and Xe abundances in the right panels of Figure 4. The correct figure is shown below.



**Figure 4.** Same as Figure 3, but for different combinations of solid and gas accretion locations. Left: predicted elemental enhancements in Jupiter’s envelope when solids and envelope are accreted at the same disk radius. Middle: predicted elemental enhancements when the gas is accreted interior to the core formation location. Right: predicted enhancements when a fraction of the solids dissolved in the enveloped originates from a core formed at 45 au and a fraction from planetesimals accreted at 5 au. Note that only scenarios where Jupiter’s core forms at 45 au fit all data. If Ar condenses out closer to the N<sub>2</sub> snowline, as suggested by recent unpublished data, core formation beyond 30 au is sufficient to explain observations.

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