

A collective description of the unusually low ratio $B_{4/2} = B(E2; 4_1^+ \rightarrow 2_1^+)/B(E2; 2_1^+ \rightarrow 0_1^+)$

TAO WANG^(a)

College of Physics, Tonghua Normal University - Tonghua 134000, PRC

received 7 February 2020; accepted in final form 16 March 2020
published online 1 April 2020

PACS 21.60.Fw – Models based on group theory
PACS 21.60.Ev – Collective models

Abstract – The unusually low value of the ratio $B_{4/2} = B(E2; 4_1^+ \rightarrow 2_1^+)/B(E2; 2_1^+ \rightarrow 0_1^+)$ along the yrast band is described in the interacting boson model-1. Two three-body interactions with $SU(3)$ symmetry are additionally introduced into the usual Hamiltonian describing the transitional behaviors from the spherical vibration to the ellipsoidal rotation. In the $SU(3)$ limit, the energy of the 4^+ state in the irrep $(2N-8, 4)$ can be lower than the one of the 4^+ state in the irrep $(2N, 0)$, which causes the anomaly.

Copyright © EPLA, 2020

Introduction. – The interacting boson model (IBM) presents an algebraic description of the nuclear structure, in which the nucleon pair is regarded as a boson [1]. In the simplest case (IBM-1), proton pair and neutron pair in an even-even nucleus are not distinguished and the total boson number N is conserved. Many collective properties of nuclei can be well described by this model with up to two-body interactions. Experience shows that, if the nucleus moves away from the closed-shell configuration, collectivity emerges and gradually dominates the behaviors of the low-lying levels.

Recently, a cluster of a few extremely neutron-deficient nuclei ^{168}Os ($N = 8$) [2], ^{166}W ($N = 9$) [3], ^{172}Pt ($N = 8$) [4] and ^{170}Os ($N=9$) [5] are experimentally found to have an unpredictably small ratio of reduced transition probabilities $B_{4/2} = B(E2; 4_1^+ \rightarrow 2_1^+)/B(E2; 2_1^+ \rightarrow 0_1^+)$. This small ratio has also been found in other nuclei away from closed shell, such as ^{50}Cr ($N = 3$) [6], ^{48}Cr ($N = 4$) [6], ^{74}Zn ($N = 4$) [7], ^{72}Zn ($N = 5$) [7], ^{114}Te ($N = 7$) [8,9] and ^{114}Xe ($N = 7$) [10]. These experimental results are very surprising. In a standard collective model, such as the geometrical collective model or IBM, the $B(E2)$ values increase with spin along the yrast band, so the ratio $B_{4/2}$ is strictly larger than unity. For an ideal rotor, this quantity is 1.43 (Alaga rule), while for a harmonic vibrator, it is 2. Although this value can become a little smaller in the IBM, it still cannot be less than 1. Despite a lot of attempts to calculate the small value of

$B_{4/2}$, such as large-scale shell models and state-of-the-art beyond-mean-field models, this anomaly cannot be reproduced so far in a convincing way [5].

One possible explanation for this anomaly is that these neutron-deficient nuclei follow seniority symmetry [11]. This pairing-correlation dominating behavior is usually expected to be found near magic neutron or proton numbers. An interesting study on the transition from non-collective seniority-like excitations to collective modes is performed with ^{206}Po ($N = 4$) and ^{204}Po ($N = 5$) recently [12]. In ref. [4] great efforts have been made to explain the $B_{4/2}$ anomaly with a seniority-conserving structure, but there are no definite results for the large model spaces involved. Furthermore, it is shown that the values of $B(E2; 2_1^+ \rightarrow 0_1^+)$ of these nuclei can be very well reproduced with existing theories. The value of $B(E2; 2_1^+ \rightarrow 0_1^+)$ is a key observable for understanding the emergence of collectivity [12] and the quantum phase transition of the shapes of atomic nuclei [13]. The $B(E2; 2_1^+ \rightarrow 0_1^+)$ values of the ^{168}Os , ^{166}W , ^{172}Pt and ^{170}Os are, respectively, 74(13) W.u., 150(9) W.u., 49(11) W.u. and 97(8) W.u., which are all larger than the value 40.60(20) W.u. of the typical γ -soft nucleus ^{196}Pt [14]. These results mean that the 2_1^+ states of these nuclei should be collective excitations. The experimental discrepancy on the small $B_{4/2}$ is merely determined by the value of the $B(E2; 4_1^+ \rightarrow 2_1^+)$ [5], which is nearly 4 times smaller than the theoretical value. If the 0_1^+ state and the 2_1^+ state belong to a different collective mode than the 4_1^+ state, a small ratio $B_{4/2}$ is possible.

^(a)E-mail: suiyueqiaoqiao@163.com

In this paper we insist that the low-lying states of these nuclei are collective, and try to resolve this $B_{4/2}$ puzzle within the framework of the IBM-1. If two three-body interactions, the $SU(3)$ third-order Casimir operator and $[\hat{L} \times \hat{Q} \times \hat{L}]^{(0)}$ (Q is the quadrupole operator in the $SU(3)$ limit and L is the angular momentum operator), are additionally introduced into the conventional Hamiltonian, it is found that this abnormally small ratio $B_{4/2}$ can be successfully reproduced. The two $SU(3)$ three-body symmetry-conserving operators have been discussed in refs. [15–20] as one of the extensions of the IBM-1. We illustrate this result with the nucleus ^{170}Os for an example. The reasons for this small $B_{4/2}$ are also clarified. In the $SU(3)$ limit, the energy of the 4^+ state in the irrep $(2N-8, 4)$ can be lower than the one of the 4^+ state in the irrep $(2N, 0)$, which is the cause of the anomaly.

Hamiltonian and numerical results. – In the IBM-1, three dynamical symmetry limits exist: the $U(5)$ limit presents quadrupole vibrations around a spherical nucleus, the $SU(3)$ limit describes an ellipsoidal nucleus having rotational structure, and the $O(6)$ limit shows a γ -soft triaxial rotation. These nuclei having a small $B_{4/2}$ value locate a transitional region from spherical vibration to prolate rotation [2–5], so the d -boson number operator in the $U(5)$ limit $\hat{C}_1[U(5)] = \hat{n}_d$, the $SU(3)$ second-order Casimir operator $\hat{C}_2[SU(3)]$ and the rotational invariant operator $\hat{C}_2[SO(3)] = \hat{L}^2$ should be considered in the Hamiltonian. However, the three terms are not enough to reproduce the anomalous result. If we insist that the small $B_{4/2}$ ratio is also a collective phenomenon, the only possible solution is that higher-order interactions should be additionally introduced into the description.

Two three-body interactions conserving $SU(3)$ symmetry are explored in this paper. One is the $SU(3)$ third-order Casimir operator $\hat{C}_3[SU(3)]$, and another is the $O(3)$ scalar shift operator $\hat{\Omega} = [\hat{L} \times \hat{Q} \times \hat{L}]^0$. These terms have been already discussed in detail in refs. [15,16], but the $SU(3)$ third-order invariant term is ignored for actual simulation. In a recent study, it is shown that a combination of the $SU(3)$ second-order and third-order Casimir operators can describe the quantum phase transition from the prolate to the oblate shapes [19], in which analytically solvable description can be obtained.

The Hamiltonian used in this paper is

$$H = \alpha \hat{n}_d + \beta \hat{C}_2[SU(3)] + \gamma \hat{C}_3[SU(3)] + \delta \hat{\Omega} + \zeta \hat{L}^2, \quad (1)$$

where $\alpha, \beta, \gamma, \delta, \zeta$ are five parameters used for fitting. We take the nucleus ^{170}Os ($N = 9$) for an example [5] to describe the small $B_{4/2}$ value for these nuclei ^{168}Os , ^{166}W , ^{172}Pt and ^{170}Os have similar level structures and $B(E2)$ features. The numerical results are performed based on our $SU(3)$ basis diagonalization program [21]. The Hamiltonian (1) is diagonalized under the $U(6) \supset SU(3) \supset SO(3)$ basis spanned by $|N(\lambda, \mu)\chi L\rangle$, where χ is the branching multiplicity occurring in the reduction of $SU(3) \downarrow SO(3)$. The basis vectors are orthonormal, so the

Table 1: The experimental and numerical results for the energy $E_{2_1^+}$ of the first 2^+ state, the energy $E_{4_1^+}$ of the first 4^+ state, the energy $E_{6_1^+}$ of the first 6^+ state, the energy $E_{8_1^+}$ of the first 8^+ state, the energy $E_{10_1^+}$ of the first 10^+ state, the reduced transitional probabilities $B(E2; 2_1^+ \rightarrow 0_1^+)$, $B(E2; 4_1^+ \rightarrow 2_1^+)$ and the ratio $B_{4/2}$ for ^{170}Os . The effective charge $e = 0.149 eb$. The experimental values are adopted from refs. [5,22].

	Expt.	Present results
$E_{2_1^+}$ (keV)	286.70(14)	283.05
$E_{4_1^+}$ (keV)	749.90(20)	729.79
$E_{6_1^+}$ (keV)	1325.42(24)	1228.43
$E_{8_1^+}$ (keV)	1945.8(4)	1946.20
$E_{10_1^+}$ (keV)	2545.2(5)	3114.18
$B(E2; 2_1^+ \rightarrow 0_1^+)(e^2b^2)$	$0.54^{+0.05}_{-0.05}$	0.540
$B(E2; 4_1^+ \rightarrow 2_1^+)(e^2b^2)$	$0.21^{+0.07}_{-0.04}$	0.204
$B_{4/2}$	0.38(11)	0.378

eigenstates of (1) can be expressed as

$$|N, L\xi; \alpha, \beta, \gamma, \delta, \zeta\rangle = \sum_{(\lambda, \mu)\chi} C_{(\lambda, \mu)\chi}^{L\xi}(\alpha, \beta, \gamma, \delta, \zeta) |N(\lambda, \mu)\chi L\rangle, \quad (2)$$

where ξ is an additional quantum number distinguishing different eigenstates with the same angular momentum L and $C_{(\lambda, \mu)\chi}^{L\xi}(\alpha, \beta, \gamma, \delta, \zeta)$ is the corresponding expansion coefficient.

Some key points for fitting the five parameters will be mentioned in the discussion part. In table 1 we list the experimental and numerical results for the energies of the 2_1^+ , 4_1^+ , 6_1^+ , 8_1^+ , 10_1^+ states and the $B_{4/2}$ value for ^{170}Os , where $\alpha = 302.40$ keV, $\beta = -30.09$ keV, $\gamma = 3.79$ keV, $\delta = -10.38$ keV, $\zeta = 18.66$ keV. It is shown that the energies of the 2_1^+ , 4_1^+ , 6_1^+ , 8_1^+ , 10_1^+ states in the yrast band and the unusually small value $B_{4/2}$ can be reproduced very well in the IBM-1 with higher-order interactions.

Discussions. – Now we reveal the reasons for this $B_{4/2}$ anomaly. The four parameters $\beta, \gamma, \delta, \zeta$ in front of the four $SU(3)$ -conserving operators are fixed, and the parameter α for the $U(5)$ limit operator changes from 0 keV to 600 keV. The evolution of the energies of the 0_1^+ , 2_1^+ , 4_1^+ , 0_2^+ , 2_2^+ and 4_2^+ states as a function of α are plotted in fig. 1. In the parameter regions, the Hamiltonian gives a γ -soft-like spectrum. The energies of the 2_2^+ state and the 0_2^+ state are lower than the one of the 4_1^+ state, which is similar to anharmonic vibrations in a spherical nucleus. The nuclei having low value $B_{4/2}$ are also located in such a γ -soft area which is adjacent to the spherical nucleus [5]. The fitting results are in line with expectations very well.

The evolution of the $B(E2)$ values of the $B(E2; 2_1^+ \rightarrow 0_1^+)$, $B(E2; 4_1^+ \rightarrow 2_1^+)$ and $B(E2; 4_2^+ \rightarrow 2_1^+)$ are plotted in

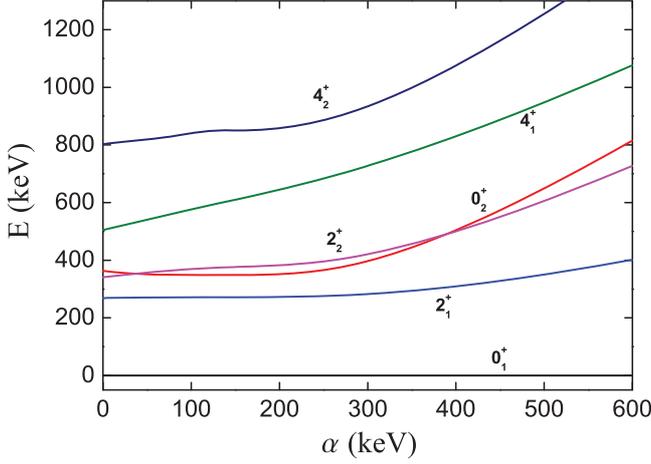


Fig. 1: The evolution of the energies of the 0_1^+ , 2_1^+ , 4_1^+ , 0_2^+ , 2_2^+ and 4_2^+ states as a function of α from 0 keV to 600 keV, here $\beta = -30.09$ keV, $\gamma = 3.79$ keV, $\delta = -10.38$ keV, $\zeta = 18.66$ keV.

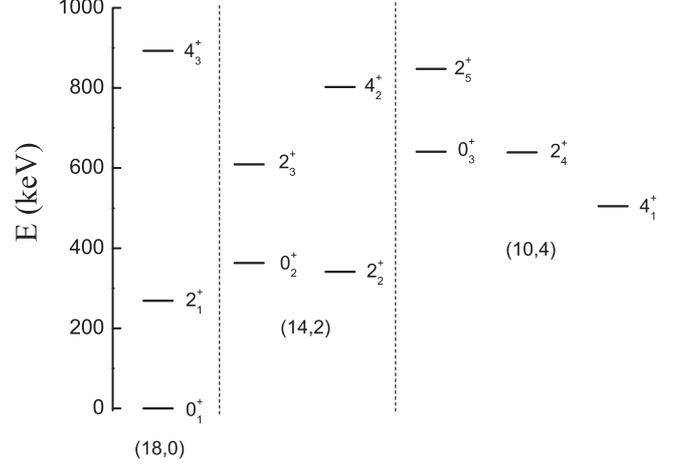


Fig. 3: The spectra of the $SU(3)$ limit with additional two three-body interactions, here $\alpha = 0$ keV, $\beta = -30.09$ keV, $\gamma = 3.79$ keV, $\delta = -10.38$ keV, $\zeta = 18.66$ keV.

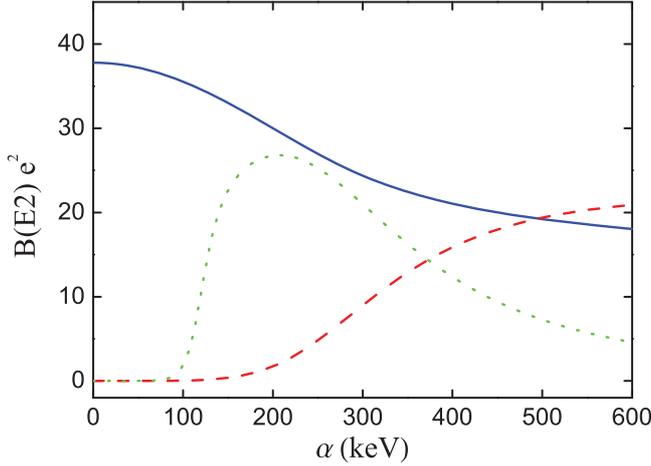


Fig. 2: The evolution of the $B(E2)$ values of the $B(E2; 2_1^+ \rightarrow 0_1^+)$ (solid blue line), $B(E2; 4_1^+ \rightarrow 2_1^+)$ (dashed red line) and $B(E2; 4_2^+ \rightarrow 2_1^+)$ (dotted green line) as a function of α from 0 keV to 600 keV, here $\beta = -30.9$ keV, $\gamma = 3.79$ keV, $\delta = -10.38$ keV, $\zeta = 18.66$ keV.

fig. 2. The $B(E2)$ operator is chosen as

$$\hat{T}(E2) = e\hat{Q}, \quad (3)$$

where e is the boson effective charge, which is usually used in IBM-1 [1,21]. The $B(E2)$ values seem very different from the results obtained from previous γ -soft spectra from the $U(5)$ limit to the $O(6)$ limit. The values of the $B(E2; 2_1^+ \rightarrow 0_1^+)$ are normal (solid blue line). When α becomes larger, the value decreases gradually, which is consistent with our experience. The values ($\alpha \leq 300$ keV) of the $B(E2; 4_1^+ \rightarrow 2_1^+)$ (dashed red line) are much smaller than the ones of the $B(E2; 2_1^+ \rightarrow 0_1^+)$ which is exactly what we want to get. When $\alpha = 0$, that is in the $SU(3)$ limit, the value of the $B(E2; 4_1^+ \rightarrow 2_1^+)$ is exactly zero, which means this transition is forbidden. The behaviors

of the values of the $B(E2; 4_2^+ \rightarrow 2_1^+)$ are also presented (dotted green line). In fig. 2, it is clear that there is a crossover point for the two evolutionary lines of the $B(E2; 4_1^+ \rightarrow 2_1^+)$ and $B(E2; 4_2^+ \rightarrow 2_1^+)$. This means 4_1^+ state and 4_2^+ state are mixed when $\alpha = 302.40$ keV.

When $\alpha = 0$, the left four interactions are $SU(3)$ symmetry conserving, which cannot break the $SU(3)$ symmetry, so the states in this limit can be labeled with the $SU(3)$ irrep (λ, μ) . Based on the results in ref. [15], the energies of the low-lying states can be directly obtained. These analytical results are the same as our numerical calculation, which can be easily checked. In fig. 3, the low-lying spectra of the $SU(3)$ limit with additional two three-body interactions are shown. It is easily seen that the energy of the 4^+ state belonging to the irrep $(10, 4)$ is lower than the one of the 4^+ state in the irrep $(18, 0)$. Thus the transition between the 4_1^+ state and 2_1^+ state is forbidden. The states belonging to different irreps (λ, μ) in the $SU(3)$ second-order and third-order Casimir operators are degenerate. The energies of the 0^+ states are decided by the two $SU(3)$ Casimir operators. The three-body interaction $[\hat{L} \times \hat{Q} \times \hat{L}]^{(0)}$ and the \hat{L}^2 interaction cannot change the energies of the 0^+ states. Thus the rotational ground-state band for the irrep $(18, 0)$ and other band structures in different irreps are generated by the $[\hat{L} \times \hat{Q} \times \hat{L}]^{(0)}$ and \hat{L}^2 interactions.

When the $U(5)$ term is added, fig. 4 predicts the partial low-lying spectra for ^{170}Os in the IBM-1 with additional two three-body interactions and table 2 presents some predicted partial absolute $B(E2)$ values for $E2$ transitions from L_i state to L_f state. This spectra resemble the ones of a γ -soft nucleus, but the $B(E2)$ modes are very different. The value of the $B(E2; 0_2^+ \rightarrow 2_1^+)$ is very small. Of particular importance, the absolute $BE(2)$ values along the yrast line really decreases with spin as expected, which further verifies the rationality of this theory. Recently the spectra with γ -soft feature having odd $B(E2)$ values have

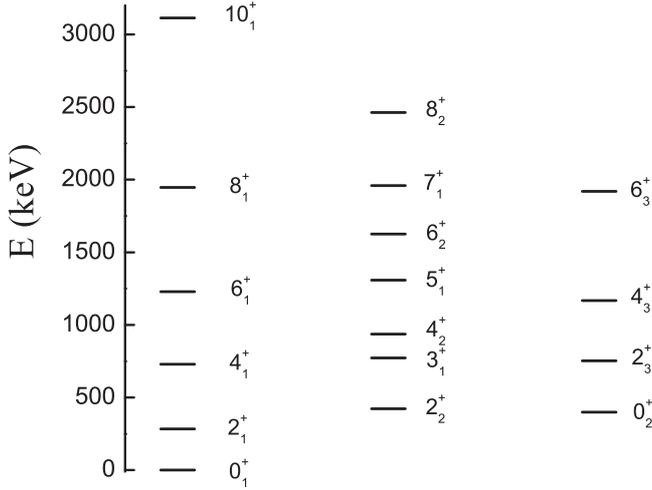


Fig. 4: Predicted partial low-lying spectra for ^{170}Os in the IBM-1 with additional two three-body interactions, here $\alpha = 302.40$ keV, $\beta = -30.09$ keV, $\gamma = 3.79$ keV, $\delta = -10.38$ keV, $\zeta = 18.66$ keV.

Table 2: Predicted partial absolute $B(E2)$ values in e^2b^2 for $E2$ transitions from L_i state to L_f state for ^{170}Os when $\alpha = 302.40$ keV, $\beta = -30.09$ keV, $\gamma = 3.79$ keV, $\delta = -10.38$ keV, $\zeta = 18.66$ keV with effective charge $e = 0.149$ eb.

L_i	L_f	Present results
2_1^+	0_1^+	0.540
4_1^+	2_1^+	0.204
6_1^+	4_1^+	0.185
8_1^+	6_1^+	0.120
10_1^+	8_1^+	0.065
2_2^+	2_1^+	0.558
0_2^+	2_1^+	0.017
4_2^+	2_1^+	0.462

attracted much attention [23]. We expect these results can be verified with future experiments.

The fitting process of the parameters needs further explanation. We start to determine the parameters from the $SU(3)$ second-order Casimir operator and the \hat{L}^2 interaction, which shows a typical ellipsoidal rotational spectra [1]. Then the $SU(3)$ third-order Casimir operator is considered [15,19]. This interaction can reduce the energies of the 0^+ excited states and the rotational bands. Next, the $[\hat{L} \times \hat{Q} \times \hat{L}]^{(0)}$ interaction is added, which can make the energy of the 4^+ state belonging to the irrep $(2N-8, 4)$ lower than the one of the 4^+ state in the irrep $(2N, 0)$. However, it is also noticed that the energies of the 6_1^+ , 8_1^+ , 10_1^+ states are lower than the one of the 4_1^+ state. Subsequently, the $U(5)$ interaction is followed. Through observing the $B_{4/2}$ value and the level ratio $E_{4_1^+}/E_{2_1^+}$, the four parameters can be preliminarily confirmed. At last, the \hat{L}^2 interaction is again considered to determine the energies of the levels along the yrast band. Final results

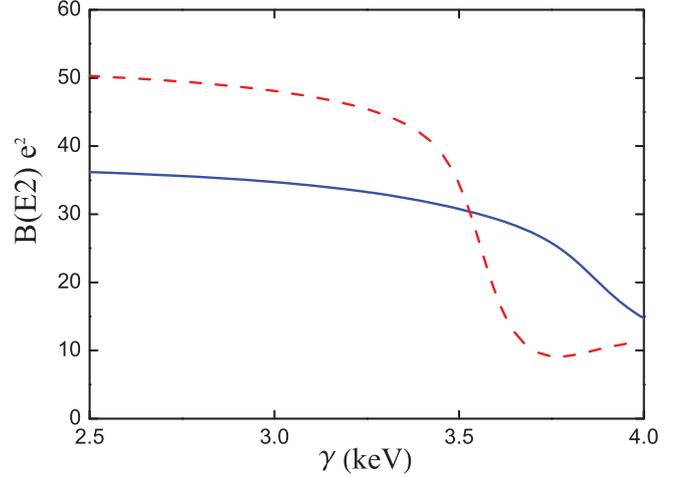


Fig. 5: The evolution of the $B(E2)$ values of the $B(E2; 2_1^+ \rightarrow 0_1^+)$ (solid blue line) and $B(E2; 4_1^+ \rightarrow 2_1^+)$ (dashed red line) as a function of γ from 2.5 keV to 4.0 keV, here $\alpha = 302.40$ keV, $\beta = -30.09$ keV, $\delta = -10.38$ keV, $\zeta = 18.66$ keV.

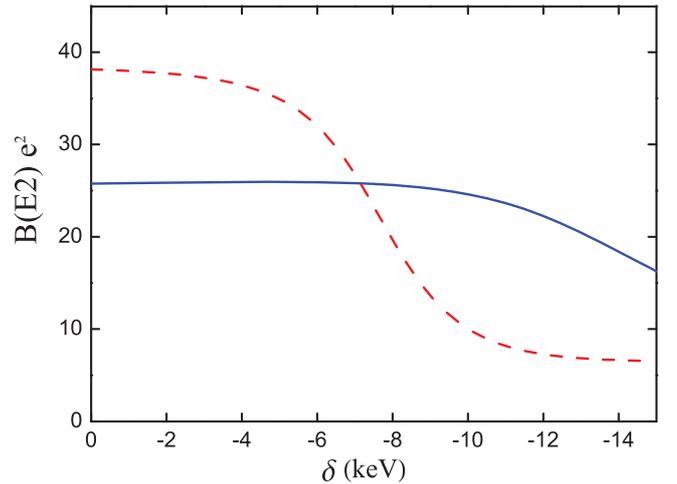


Fig. 6: The evolution of the $B(E2)$ values of the $B(E2; 2_1^+ \rightarrow 0_1^+)$ (solid blue line) and $B(E2; 4_1^+ \rightarrow 2_1^+)$ (dashed red line) as a function of δ from 0 keV to -15 keV, here $\alpha = 302.40$ keV, $\beta = -30.09$ keV, $\gamma = 3.79$ keV, $\zeta = 18.66$ keV.

need to be adjusted and validated repeatedly until they match the experimental data.

The $B_{4/2}$ anomaly occurs in a special parameter region. The evolution of the $B(E2)$ values of the $B(E2; 2_1^+ \rightarrow 0_1^+)$ and $B(E2; 4_1^+ \rightarrow 2_1^+)$ as a function of γ from 2.5 keV to 4 keV are plotted in fig. 5. It can be seen that even the γ value becomes a little smaller, such as $\gamma = 3.40$ keV, the anomaly disappears. This result may be contrary to our experience for the two three-body interactions are also large. In previous calculation, the higher-order interactions are not so important for the Hamiltonian with up to two-body interactions can also give the main results observed experimentally. Similar situations also hold for the δ and β parameters, see fig. 6 and fig. 7.

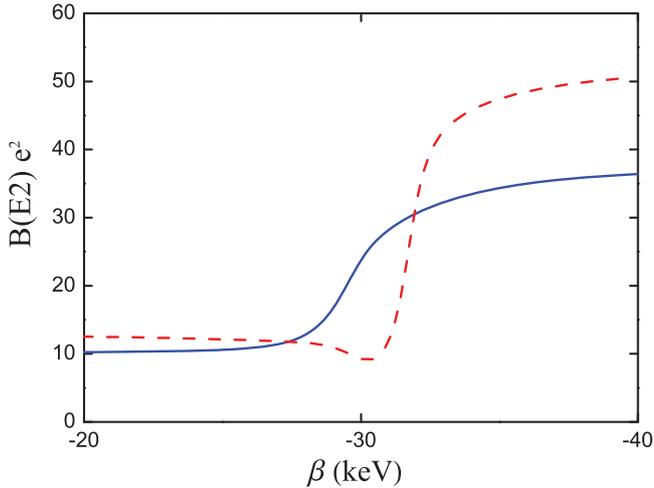


Fig. 7: The evolution of the $B(E2)$ values of the $B(E2; 2_1^+ \rightarrow 0_1^+)$ (solid blue line) and $B(E2; 4_1^+ \rightarrow 2_1^+)$ (dashed red line) as a function of β from -20 keV to -40 keV, here $\alpha = 302.40$ keV, $\gamma = 3.79$ keV, $\delta = -10.38$ keV, $\zeta = 18.66$ keV.

We expect the small $B_{4/2}$ value can be also reproduced in the shell model with adding appropriate three-body interactions. The microscopic shell model foundation of the IBM is still under investigation, and the reason of the emergence of the three-body interactions is out of our knowledge. It is essential to investigate the small value $B_{4/2}$ in the IBM-2, in which the proton pair and the neutron pair are treated separately. IBM-2 has a direct correspondence to the shell model. If we can confirm the function of the two three-body interactions in the IBM-2, some clues of the three-body interactions added in the shell model may be obtained. Another method to construct the three-body interactions is to explore the $SU(3)$ shell model [24,25]. Similar to our approach, the low 4_1^+ state having forbidden transition to the 2_1^+ state may be generated [20].

It should be noticed that the abruptly decreasing of the $B(E2)$ values can also occur for the $B(E2; 6_1^+ \rightarrow 4_1^+)$ or even the $B(E2; 2_1^+ \rightarrow 0_1^+)$ for appropriate parameters. The small value $B(E2; 6_1^+ \rightarrow 4_1^+)$ compared to the value $B(E2; 4_1^+ \rightarrow 2_1^+)$ can be found in ^{72}Zn ($N = 5$). The previous value is $134_{-31}^{+57} \text{ e}^2\text{fm}^4$ and the latter one is $361_{-47}^{+57} \text{ e}^2\text{fm}^4$ [7]. For $^{70,72,74}\text{Zn}$, our theory may provide a self-consistent result, this will be discussed in future work. An experimental measurement of the value $B(E2; 6_1^+ \rightarrow 4_1^+)$ in ^{168}W ($N = 10$) is highly anticipated [3].

Conclusions. – In this paper, the $B(E2)$ anomaly within the yrast band is resolved in the collective framework of the IBM-1. Two $SU(3)$ -conserving three-body interactions are additionally introduced into the conventional Hamiltonian describing the transitional behaviors from the spherical vibration to the ellipsoidal rotation. A γ -soft-like spectra is generated with the small $B_{4/2}$ value. These results are very different from our common

experience, just like the experimental anomaly is contrary to our expectation. In the IBM-1, this is the only convincing way to reproduce the anomalous small $B_{4/2}$ value, thus we believe that the three-body interactions may be more important than we thought before, which should be a necessary ingredient to describe the collective nature of nuclei. A complete study in the IBM-2 to discuss the $B_{4/2}$ anomaly will be done in future work, and a similar realization of our work in the shell model, especially the $SU(3)$ shell model is highly expected.

REFERENCES

- [1] IACHELLO F. and ARIMA A., *The Interacting Boson Model* (Cambridge University Press) 1987.
- [2] GRAHN T., STOLZE S., JOSS D. T., PAGE R. D., SAYGI B., O'DONNELL D., AKMALI M., ANDGREN K., BIANCO L., CULLEN D. M., DEWALD A., GREENLEES P. T., HEYDE K., IWASAKI H., JAKOBSSON U., JONES P., JUDSON D. S., JULIN R., JUUTINEN S., KETELHUT S., LEINO M., LUMLEY N., MASON P. J. R., MÖLLER O., NOMURA K., NYMAN M., PETTS A., PEURA P., PIETRALLA N., PISSULLA T., RAHKILA P., SAPPLE P. J., SARÉN J., SCHOLEY C., SIMPSON J., SORRI J., STEVENSON P. D., UUSITALO J., WATKINS H. V. and WOOD J. L., *Phys. Rev. C*, **94** (2016) 044327.
- [3] SAYGI B., JOSS D. T., PAGE R. D., GRAHN T., SIMPSON J., O'DONNELL D., ALHARSHAN G., AURANEN K., BÄCK T., BOENING S., BRAUNROTH T., CARROLL R. J., CEDERWALL B., CULLEN D. M., DEWALD A., DONCEL M., DONOSA L., DRUMMOND M. C., ERTUGRAL F., ERTRK S., FRANSEN C., GREENLEES P. T., HACKSTEIN M., HAUSCHILD K., HERZAN A., JAKOBSSON U., JONES P. M., JULIN R., JUUTINEN S., KONKI J., KRÖLL T., LABICHE M., LOPEZ-MARTENS A., MCPPEAKE C. G., MORADI F., MÖLLER O., MUSTAFA M., NIEMINEN P., PAKARINEN J., PARTANEN J., PEURA P., PROCTER M., RAHKILA P., ROTHER W., RUOTSALAINEN P., SANDZELIUS M., SARÉN J., SCHOLEY C., SORRI J., STOLZE S., TAYLOR M. J., THORNTHWAITTE A. and UUSITALO J., *Phys. Rev. C*, **96** (2017) 021301.
- [4] CEDERWALL B., DONCEL M., AKTAS Ö., ERTOPRAK A., LIOTTA R., QI C., GRAHN T., CULLEN D. M., NARA SINGH B. S., HODGE D., GILES M., STOLZE S., BADRAN H., BRAUNROTH T., CALVERLEY T., COX D. M., FANG Y. D., GREENLEES P. T., HILTON J., IDEGUCHI E., JULIN R., JUUTINEN S., RAJU M. K., LI H., LIU H., MATTÄ S., MODAMIO V., PAKARINEN J., PAPADAKIS P., PARTANEN J., PETRACHE C. M., RAHKILA P., RUOTSALAINEN P., SANDZELIUS M., SARÉN J., SCHOLEY C., SORRI J., SUBRAMANIAM P., TAYLOR M. J., UUSITALO J. and VALIENTE-DOBÓN J. J., *Phys. Rev. Lett.*, **121** (2018) 022502.
- [5] GOASDUFF A., LJUNGVALL J., RODRÍGUEZ T. R., BELLO GARROTE F. L., ETILE A., GEORGIEV G., GIACOPPO F., GREUTE L., KLINTEFJORD M., KUSOĞLU A., MATEA I., ROCCIA S., SALSAC M.-D. and SOTTY C., *Phys. Rev. C*, **100** (2019) 034302.
- [6] HERTZ-KINTISH D., ZAMICK L. and ROBINSON S. J. Q., *Phys. Rev. C*, **90** (2014) 034307.

- [7] LOUCHART C., OBERTELLI A., GÖRGEN A., KORTEN W., BAZZACCO D., BIRKENBACH B., BRUYNEEL B., CLÉMENT E., COLEMAN-SMITH P. J., CORRADI L., CURIEN D., DE ANGELIS G., DE FRANCE G., DELAROCHE J.-P., DEWALD A., DIDIERJEAN F., DONCEL M., DUCHÊNE G., EBERTH J., ERDURAN M. N., FARNEA E., FINCK C., FIORETTO E., FRANSEN C., GADEA A., GIROD M., GOTTARDO A., GREBOSZ J., HABERMANN T., HACKSTEIN M., HUYUK T., JOLIE J., JUDSON D., JUNGCLAUS A., KARKOUR N., KLUPP S., KRÜCKEN R., KUSOGLU A., LENZI S. M., LIBERT J., LJUNGVALL J., LUNARDI S., MARON G., MENEGAZZO R., MENGONI D., MICHELAGNOLI C., MILLION B., MOLINI P., MÖLLER O., MONTAGNOLI G., MONTANARI D., NAPOLI D. R., ORLANDI R., POLLAROLO G., PRIETO A., PULLIA A., QUINTANA B., RECCHIA F., REITER P., ROSSO D., ROTHER W., SAHIN E., SALSAC M.-D., SCARLASSARA F., SCHLARB M., SIEM S., SINGH P. P., SÖDERSTRÖM P.-A., STEFANINI A. M., STÉZOWSKI O., SULIGNANO B., SZILNER S., THEISEN CH., UR C. A., VALIENTE-DOBÓN J. J. and ZIELINSKA M., *Phys. Rev. C*, **87** (2013) 054302.
- [8] CAKIRLI R. B., CASTEN R. F., JOLIE J. and WARR N., *Phys. Rev. C*, **70** (2004) 047302.
- [9] MÖLLER O., WARR N., JOLIE J., DEWALD A., FITZLER A., LINNEMANN A., ZELL K. O., GARRETT P. E. and YATES S. W., *Phys. Rev. C*, **71** (2005) 064324.
- [10] DE ANGELIS G., GADEA A., FARNEA E., ISOCRATE R., PETKOV P., MARGINEAN N., NAPOLI D. R., DEWALD A., BELLATO M., BRACCO A., CAMERA F., CURIEN D., POLI M. D., FIORETTO E., FITZLER A., KASEMANN S., KINTZ N., KLUG T., LENZI S., LUNARDI S., MENEGAZZO R., PAVAN P., PEDROZA J. L., PUCKNELL V., RING C., SAMPSON J. and WYSS R., *Phys. Lett. B*, **535** (2002) 93.
- [11] DE SHALIT A. and TALMI I., *Nuclear Shell Theory* (Academic, New York) 1963.
- [12] STOYANOVA M., RAINOVSKI G., JOLIE J., PIETRALLA N., BLAZHEV A., BECKERS M., DEWALD A., DJONGOLOV M., ESMAYLZADEH A., FRANSEN C., GERHARD L. M., GLADNISHKI K. A., HERB S., JOHN P. R., KARAYONCHEV V., KEATINGS J. M., KERN R., KNAFLA L., KOICHEVA D., KORNWEBEL L., KRÖLL TH., LEY M., MASHTAKOV K. M., MÜLLER-GATERMANN C., RÉGIS J.-M., SCHECK M., SCHOMACKER K., SINCLAIR J., SPAGNOLETTI P., SÜRDER C., WARR N., WERNER V. and WIEDERHOLD J., *Phys. Rev. C*, **100** (2019) 064304.
- [13] CEJNAR P., JOLIE J. and CASTEN R. F., *Rev. Mod. Phys.*, **82** (2010) 2155.
- [14] HUANG X. L., *Nucl. Data Sheets*, **108** (2007) 1093.
- [15] VANDEN BERGHE G., DE MEYER H. E. and VAN ISACKER P., *Phys. Rev. C*, **32** (1985) 1049.
- [16] VANTHOURNOUT J., *Phys. Rev. C*, **41** (1990) 2380.
- [17] SMIRNOV Y. F., SMIRNOVA N. A. and VAN ISACKER P., *Phys. Rev. C*, **61** (2000) 041302(R).
- [18] FORTUNATO L., ALONSO C. E., ARIAS J. M., GARCÍA-RAMOS J. E. and VITTURI A., *Phys. Rev. C*, **84** (2011) 014326.
- [19] ZHANG Y., PAN F., LIU Y. X., LUO Y. A. and DRAAYER J. P., *Phys. Rev. C*, **85** (2012) 064312.
- [20] ZHANG Y., PAN F., DAI L. R. and DRAAYER J. P., *Phys. Rev. C*, **90** (2014) 044310.
- [21] PAN F., WANG T., HUO Y. S. and DRAAYER J. P., *J. Phys. G: Nucl. Part. Phys.*, **35** (2008) 1263.
- [22] BAGLIN C. M., MCCUTCHAN E. A., BASUNIA S. and BROWNE E., *Nucl. Data Sheets*, **153** (2018) 1.
- [23] GARRETT P. E., WOOD J. L. and YATES S. W., *Phys. Scr.*, **93** (2018) 063001.
- [24] ELLIOTT J. P., *Proc. R. Soc. London, Ser. A*, **245** (1958) 128.
- [25] ELLIOTT J. P. and HARVEY M., *Proc. R. Soc. London, Ser. A*, **272** (1963) 557.